



Proposed Single Open Reflector for the GALILEO Mission

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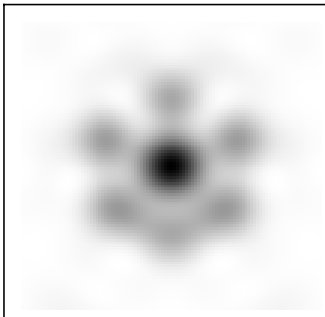
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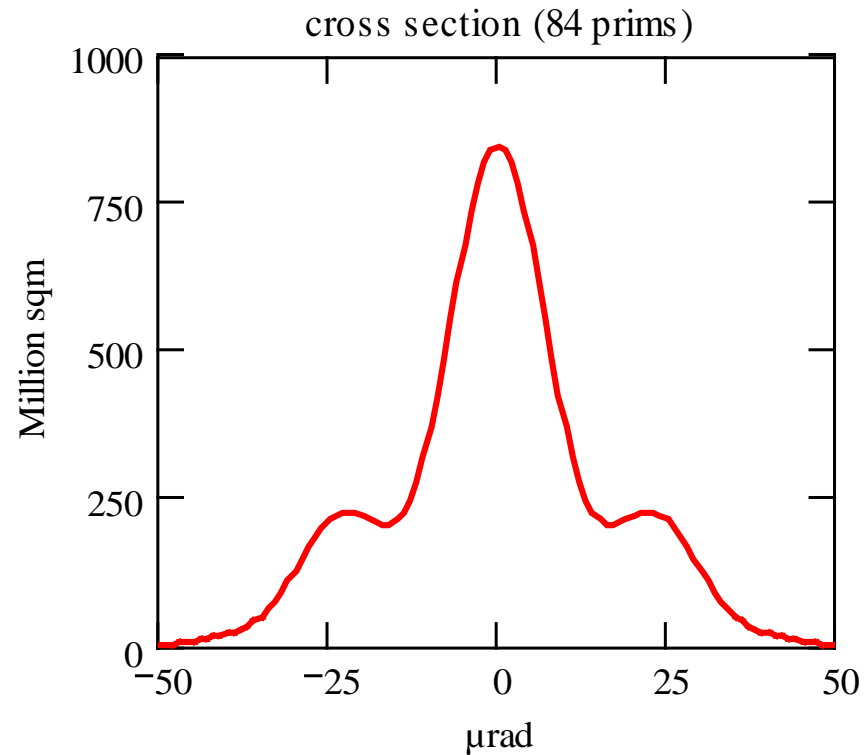
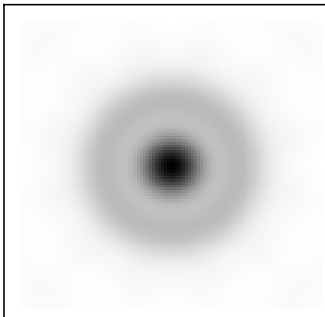
Far Field of the GALILEO Prism Array

velocity aberration 23.7...24.2 μrad

Single prism



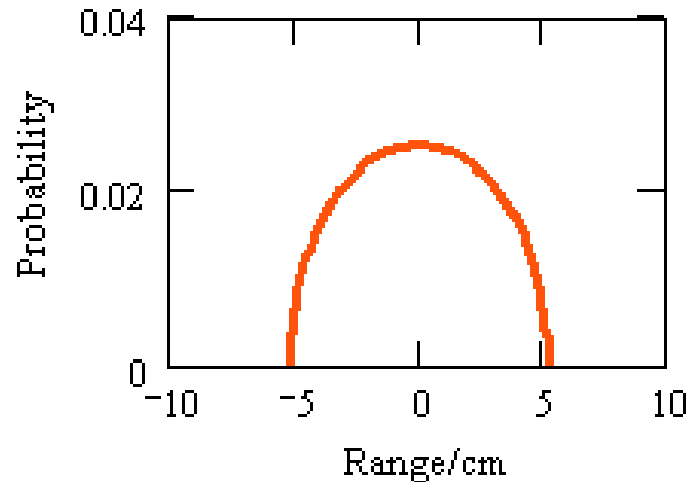
Average





Expected Signature of the GALILEO Array

circular array, radius $R=25\text{cm}$, angle of incidence: $\alpha=12^\circ$



Single shot precision:

$$\sigma = \frac{R}{2} \cdot \sin(\alpha)$$

$$\sigma = 2.6\text{cm} \quad (12^\circ)$$

Normal point precision: $\sigma_{\text{res}} = \frac{\sigma}{\sqrt{N}}$

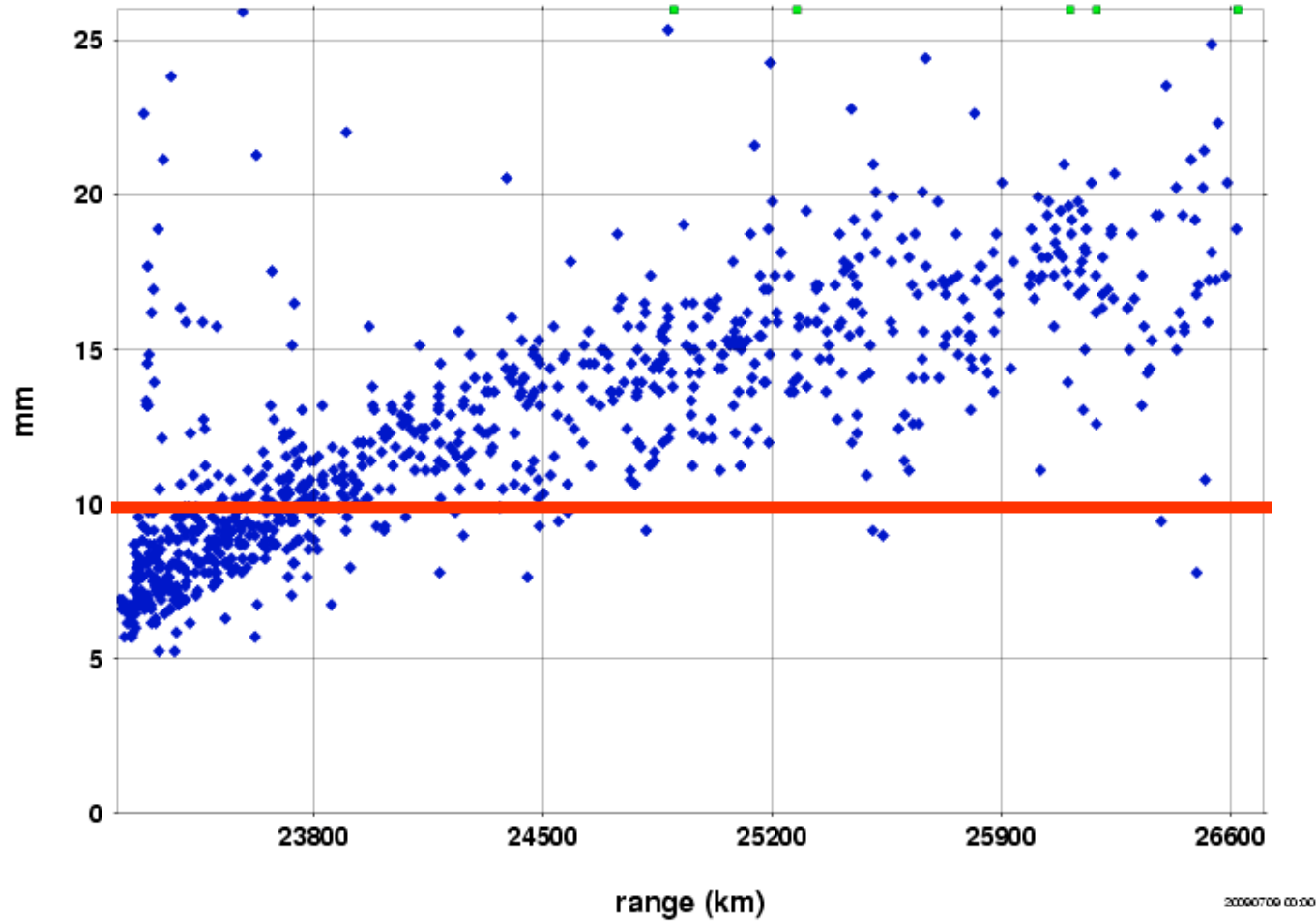
25 measurements required to get 5mm precision



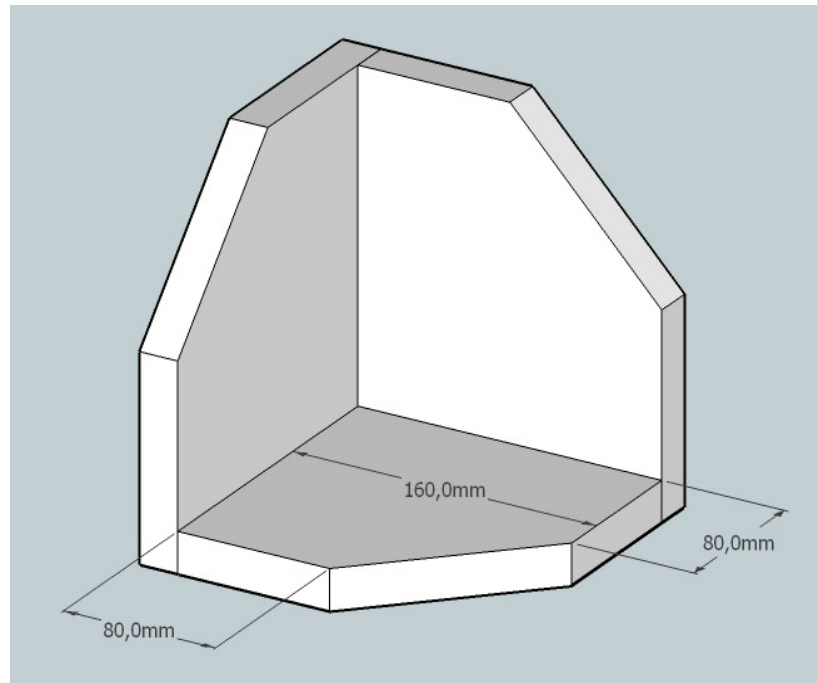
Zimmerwald, Switzerland 7810

GIOVE-B normal point rms, from 20080701 through 20090630

ave 12.51 ± 4.28 max 36.12 min 5.25 for 881 data points



The Proposed Open Reflector Geometry

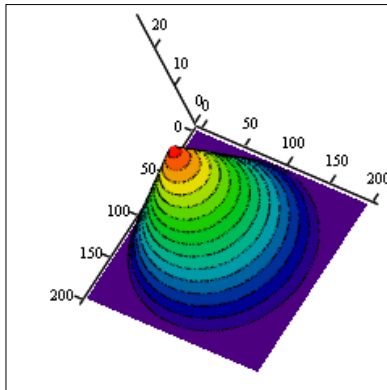


Main Features of the Reflector:

- no pulse spreading
- high effective radar cross section
- small (up to 20cm diam. front face)
- low weight

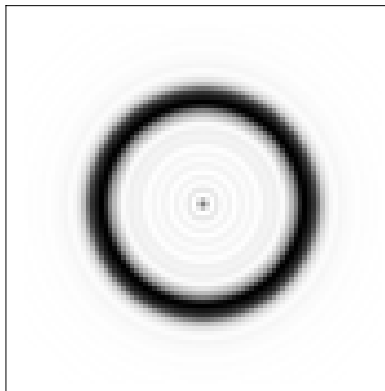
Far Field of the Open Reflector with Conical Wave Front

Wave front

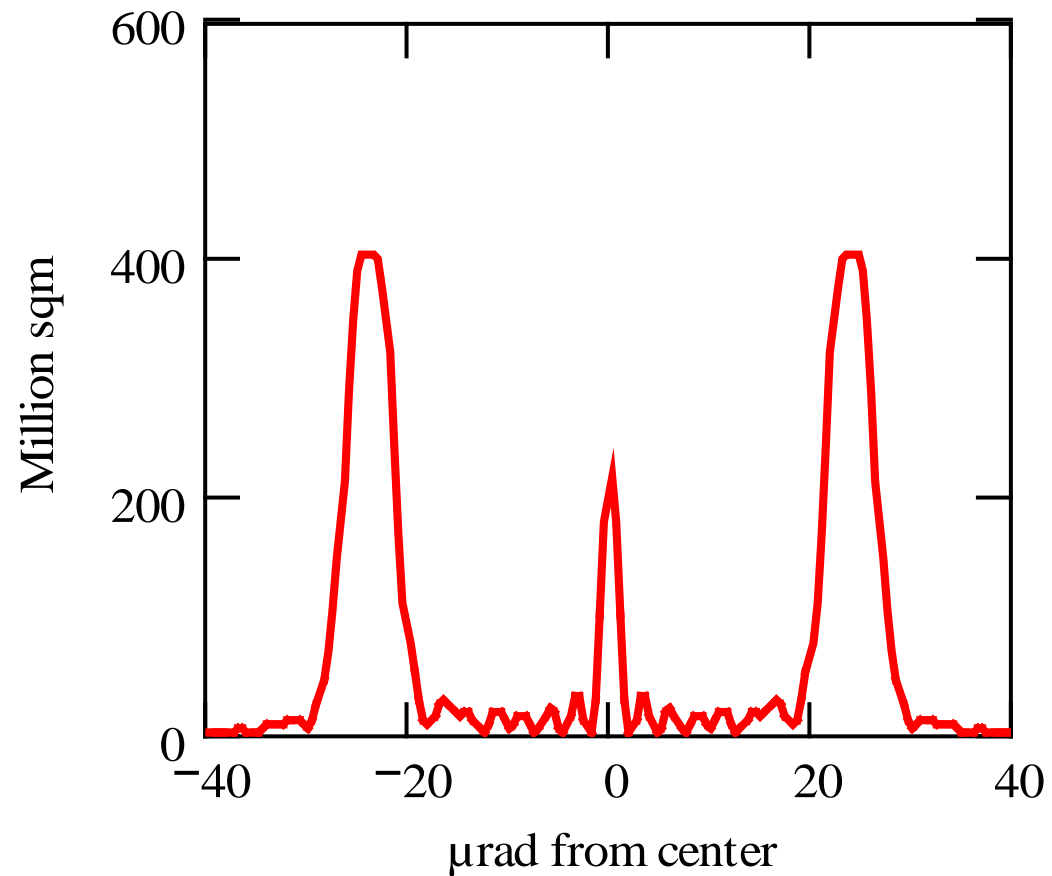


F

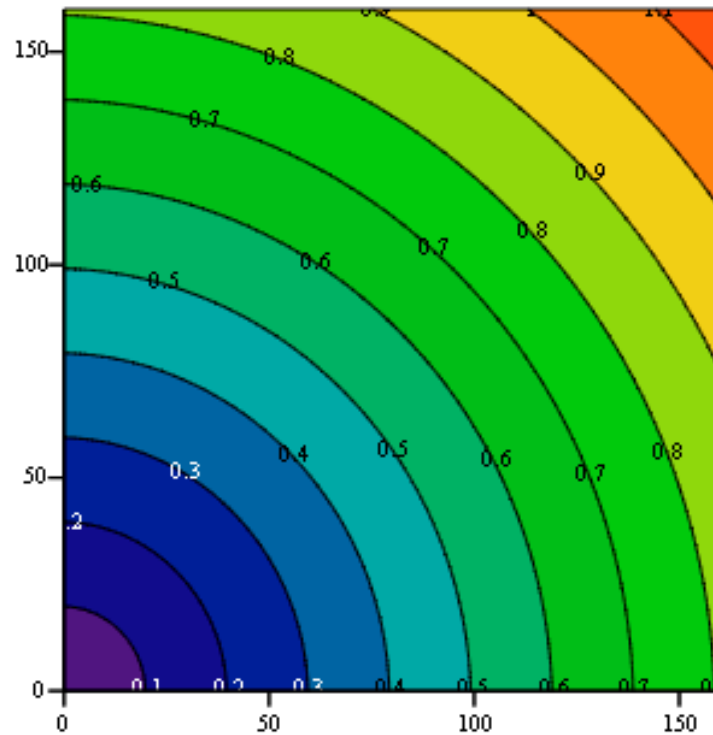
Far Field



cross section

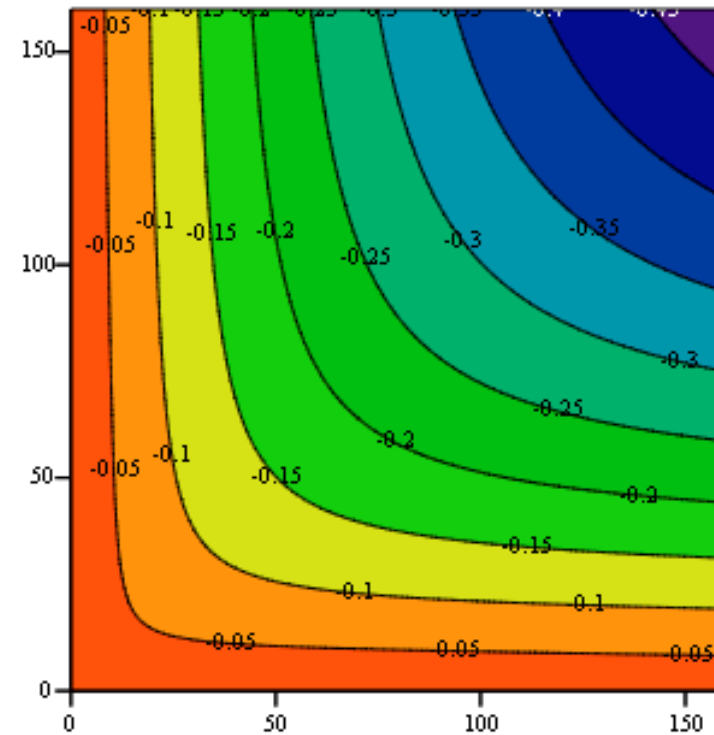


Conical Shape of the Mirrors



A

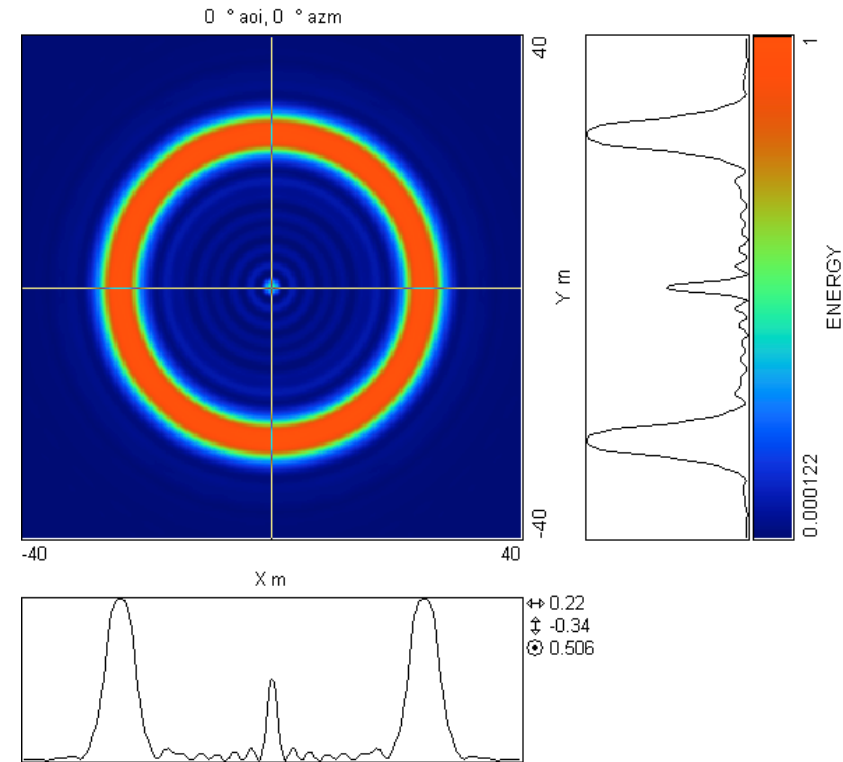
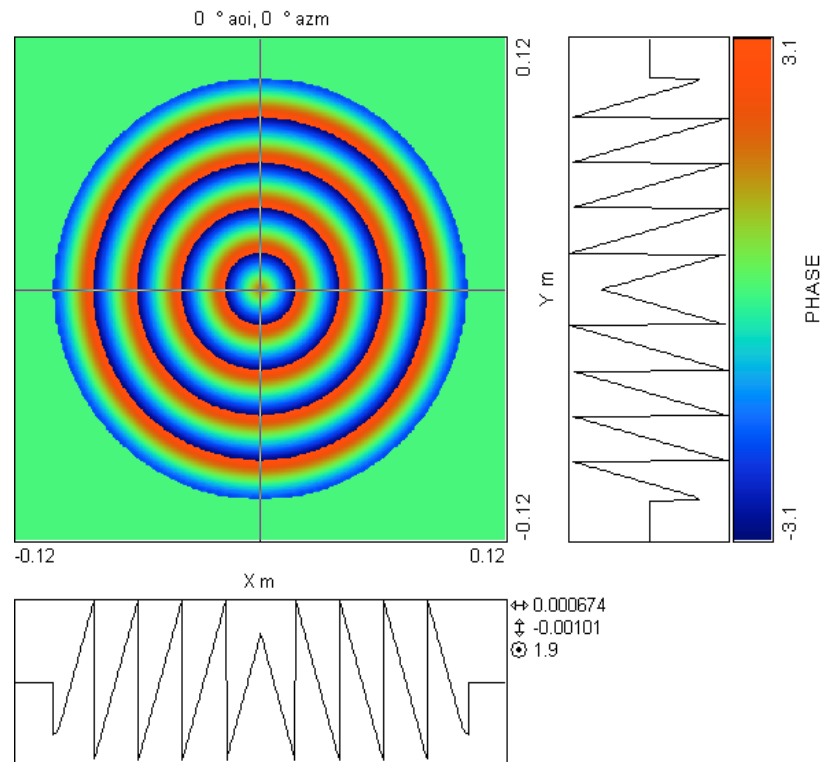
difference to perfect LRR
in μm



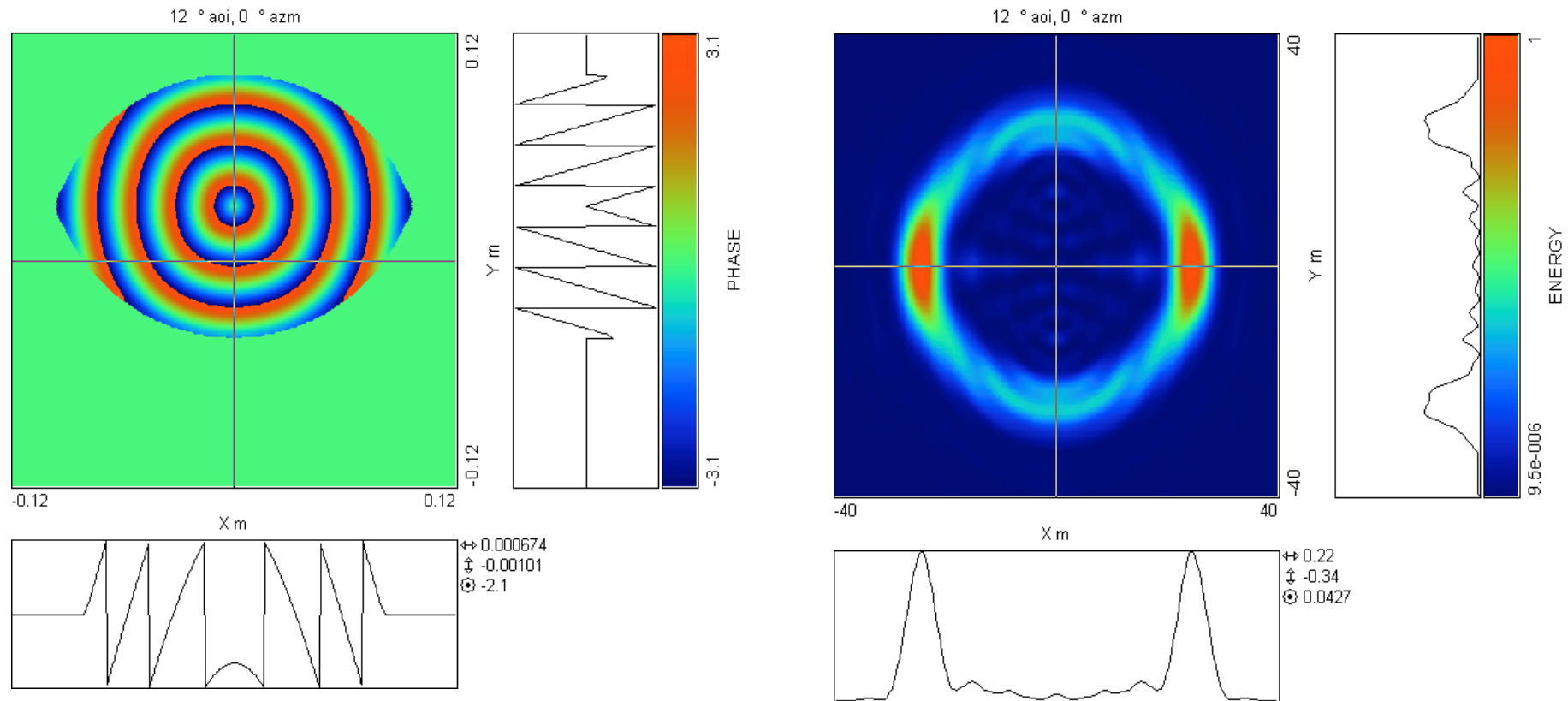
A - B

remaining deformation after
introducing 2.1 arcsec offsets

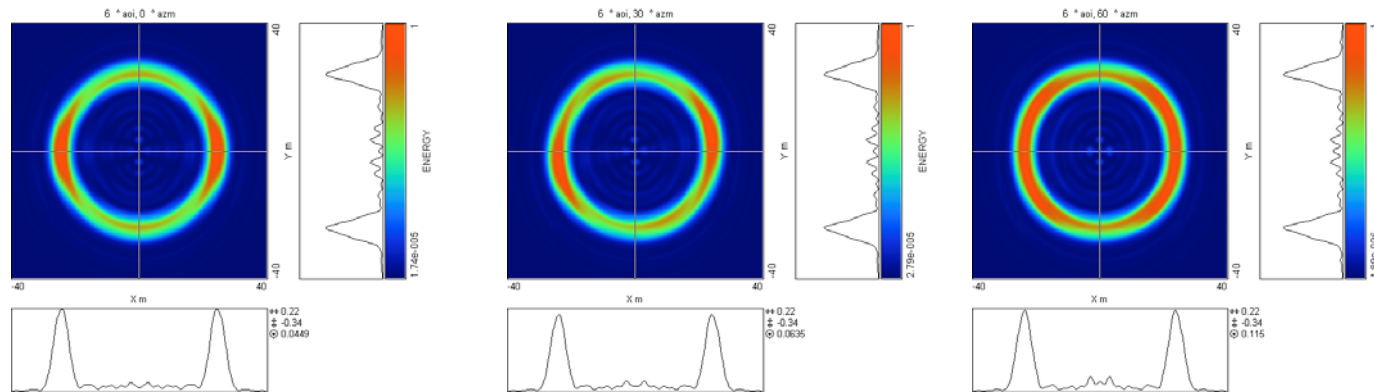
Wave Front and Far Field from ASAP, $\alpha=0^\circ$



Wave Front and Far Field from ASAP, $\alpha=12^\circ$



Far Field: Dependence on Orientation



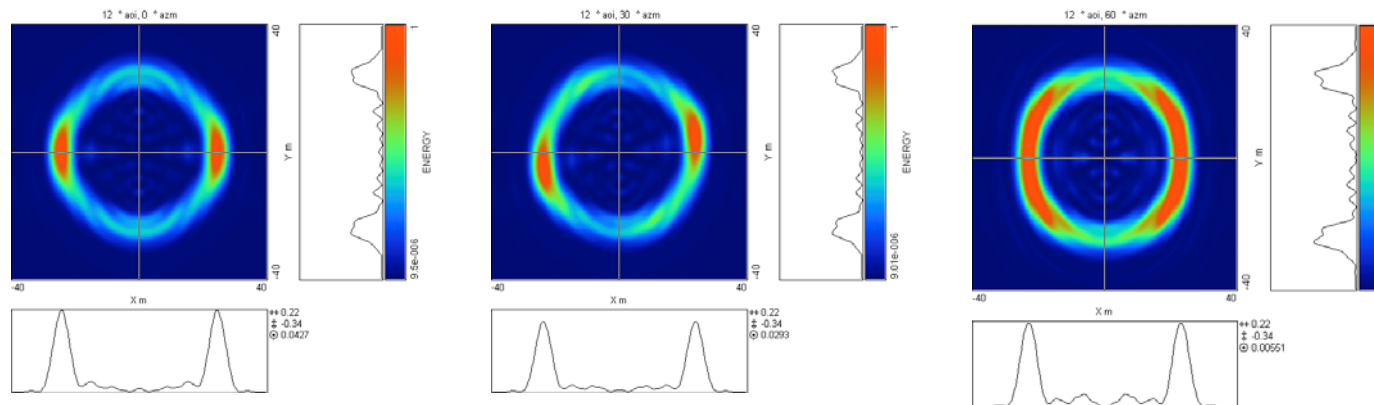
$\alpha=6^\circ$

0°

30°

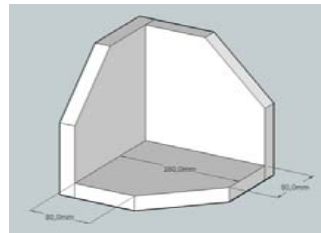
60°

Azimuth



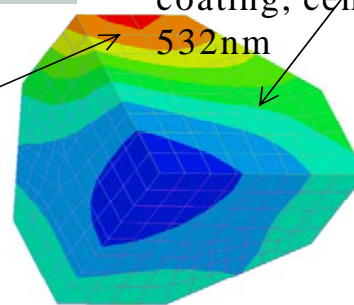
$\alpha=12^\circ$

Thermal Simulations of Orbital Heat Loads

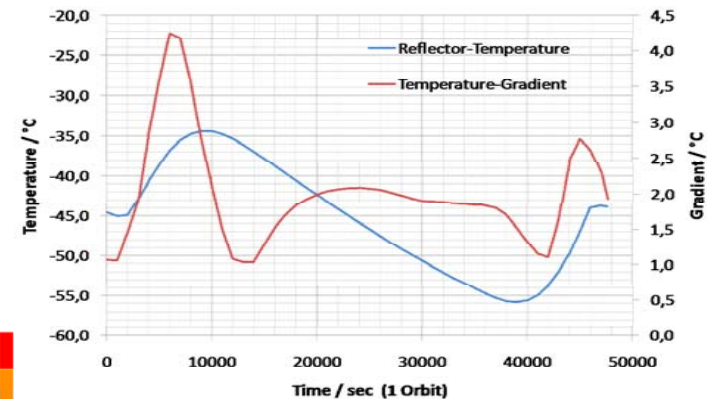


Zerodur mirror, high-reflective dielectrical coating, centered @ 532nm

Partial Sun Heating shortly after Eclipse



Color-coded temperature plot,
Red = -35°C
Blue = -39°C



Preliminary Results, Design will be further optimized

- High transient variation of the average temperature is approx. 20°C (blue line)
- Low spatial variation of approx. 4.5°C at 6000sec shortly after eclipse phase from 45967sec to 1619sec (red line)



Conclusion

The proposed open reflector is advantageous compared to the standard prism array because of:

- no pulse spreading at all angles
- equal or higher effective cross section but smaller dimension (higher antenna gain)
- at low elevation (high angle of incidence) a single received photon is needed only to achieve the full resolution of the SLR system

Main problem area:

thermal distortion caused by sun heating should be kept small by the use of selected Zerodur substrate, dielectric coating and sun shield