21st International Workshop on Laser Ranging «Laser Ranging for Sustainable Millimeter Geoscience» Canberra, Australia, November 05 – 09, 2018

Experimental determination of photometric characteristics of the BLITS-M satellite

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BLITS-M satellite

2

The BLITS-M spacecraft is a retroreflector spherical system for the fundamental geodesic support of the GLONASS system

BLITS-M consists of a spherical lens and two meniscus, one of which has an interference mirror coating.







Far-field diffraction pattern of BLITS-M

Ray tracing







Calculation of star magnitude taking into account orbital altitude (850 km and 1500 km)





Measurements of the star magnitude of the BLITS at the AOLC on 02/20/2013

Star magnitude of BLITS



Time



Measurement scheme of star magnitude



ИСИ – solar simulator; ПИ – radiation receiver (luxmeter); φ – phase angle; α – the angle of rotation of the object, defined between the normal to the interface and the direction to the ИСИ.

Results of laboratory measurements of star magnitude of BLITS

Star magnitude



Results of laboratory measurements of star magnitude of **BLITS-M**

Star magnitude



Zenith angle $\theta_z = 0^\circ$





- 1. The star magnitude at the point of observation, due to the reflection of solar radiation from the hemisphere with an interference mirror coating, does not depend on the phase angle and the rotation of the satellite around its axis. The range of magnitude, depending on the zenith angle, is from 10,5^{*m*} to 11,5^{*m*}.
- 2. The star magnitude at the observation point, due to the incidence of solar radiation on the transmissive (uncoated) hemisphere, depends on the phase angle and the rotation angle of the satellite around its axis. The magnitude in this case varies from 10,8^{*m*} to 12,6^{*m*} at zenith angle $z = 0^{\circ}$ and from 11,8^{*m*} to 13,8^{*m*} at $z = 60^{\circ}$.

As a result, we expect that there will be no major problems with the detection of the BLITS-M.



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Thank you for your attention!



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