

The Impact of Single Photon SLR Technology on Large Scale Topo-Bathymetric Mapping

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Why Single Photon Lidars ?

- Most efficient 3D lidar imager possible; each range measurement requires only one detected photon as opposed to hundreds or thousands in conventional laser pulse time of flight (TOF) altimeters and imagers
- High efficiency translates to either
 - significantly less mass, volume, and prime power ; or
 - orders of magnitude more imaging capability
- Single photon sensitivity, combined with fast recovery multistop timing receivers, enables the lidar to operate in daylight and penetrate porous obscurations such as vegetation, ground fog, thin clouds, water columns, camouflage, etc.
- Makes contiguous, high resolution topographic and bathymetric mapping on a single overflight possible with very modest laser powers and telescope apertures – even from orbital altitudes.
- SPLs derive much of their heritage from early NASA work on the single photon sensitive SLR2000 system and the "Microaltimeter".



2nd Generation "Leafcutter"

4" Scanner

• Transmitter is a low-energy (6 μ J), high rep-rate (to 22 kHz), frequency doubled (532 nm), passively Q-switched microchip laser with a 710 psec FWHM pulsewidth.

•Diffractive Optical Element (DOE) splits green output into 100 beamlets (~50 nJ @ 20 kHz = 1 mW per beamlet) in a 10 x 10 array. Residual 1064 nm energy can be used for polarimetry.

• Returns from individual beamlets are imaged by a 3 inch diameter telescope onto matching anodes of a 10x10 segmented anode microchannel plate photomultiplier.

•Each anode output is input to one channel of a 100 channel multi-stop timer to form a 100 pixel 3D image on each pulse. Individual images are <u>contiguously</u> mosaiced together via the aircraft motion and an optical scanner (100 pixels @ 22 kHz = 2.2 million 3D pixels/sec!).

• The high speed, 4" aperture, dual wedge scanner can generate a wide variety of patterns. The transmitter and receiver share a common telescope and scanner with matching (small) FOV for solar noise reduction.

Lidar mounted on camera tripod for rooftop testing

3" Telescope

Data

Recorder

Optical

Bench



Sample Data from "Leafcutter"

Single Overflight at AGLs between 2 kft (left) and 8.2 kft (right)





Sigma Single Photon LiDARs (SPLs) **Operating altitudes ranging from 2,000 to 65,000 ft.**















Multiple Altimeter Beam Experimental LiDAR (MABEL) - Pushbroom Nominal Flight AGL: 65,000 ft Platform : NASA ER-2 Customer: NASA GSEC 16 beams @ 10 kHz = 0.16 Megapixels/sec High Altitude LiDAR (HAL) -Scanning Nominal Flight AGL:: 25,000 to 36,000 ft **Platform : Various** Customer: Government Agencies 100 beams@ 32 kHz = 3.2 Megapixels/sec

High Resolution Quantum LiDAR System (HRQLS1 and 2) - Scanning Flight Altitude: 6,500 to 15,000 ft AGL Platform : King Air B200 Sigma Self-funded 100 beams @ 25 kHz/60 khz = 2.5/6 Mp/sec Miniature Airborne Topographic Mapper (Mini-ATM) - Scanning Flight Altitude: 2000 to 6000 ft AGL Platform : Viking 300 UAV **Customer: NASA Wallops**



Conventional Lidar vs SPL





HRQLS 3D Mapping, Easton, MD

Sample: Single Pass AGL: 7500 ft Velocity: 180 knots

Profile: 110 m long 10 m height









Swatantran, A; Tang, H.; Barrett, T.; DeCola, P.,; Dubayah, R. "Rapid, High-Resolution Forest Structure and Terrain Mapping over Large Areas using Single Photon Lidar", Scientific Reports 6 (2016) HRQLS-1 SPL point cloud profiles showing different growth patterns within a 1 square kilometer of forested area in Garrett County, MD.(a) Short even aged stand with little understory vegetation. (b) aged Uneven stand composed of tall trees and dense midstory vegetation. (c) Even aged stand with some mid and understory growth. (d) Tall open stand with

distinct understory vegetation (Courtesy of the University of Maryland [8]).



SPL 3D Mapping: Garrett County, MD



Garrett County, MD (1,700 km²) acquired in:

- 12 hours (including ferry and turns)
- 1.4 Km swaths, 180 Knots
- 50% overlap
- Mean Density: 12 points/m²

RGB Overlay of 3D Lidar Data





Garrett County Coal Mine





Velocity = 150 knot (278 km/hr) Scan Angle = <u>+</u>17 deg AGL = 7500 ft (2.3 km) Swath = 1.36 km Coverage: 378 km²/hr

Coal Piles



SPL 3D Mapping: Monterrey, CA

EO

Naval Postgraduate School Campus

Lidar



Single Pass, Altitude 7500 ft, 180 kts,1.4 km swath, >12 points/m²



Classified Point Cloud, colored by RGB





SPL Vertical Accuracy Assessment: Ground Truth Points



Ground Truth: 21 points measured by the Naval Postgraduate School to 3cm vertical accuracy and supplied in the same coordinate system as the point cloud (WGS84, UTM Zone 10N, ellipsoid height, meter) Standard deviation relative to ground truth: 9.3cm



SPL Bathymetry, Pt. Lobos, CA

Sample: Double Pass AGL: 7500 ft Velocity: 180 knots



Profile

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HRQLS-1 High Density Images

HRQLS-1 image of a Cruise ship docked in Ft. Lauderdale, Florida

HRQLS-1 image of a powerline grid in North Carolina





Multistage Noise Filter

HRQLS data over town of Oakland, MD

Originally, our SPL data was filtered following download of all surface data and noise into a point cloud generator. We have recently developed and successfully tested a multistage filter that, due to the high spatial correlation of up to 100 surface returns within a single pulse, can be implemented on a single pulse or small number of consecutive pulses. The goal is to reduce both onboard data storage requirements and data download/preprocessing times and ultimately permit the real time display or transmission of "clean" 3D images.

- 1st stage rapidly isolates the surface returns (including forested areas) from the solar background and typically discards ~ 94% of the raw noise
- 2nd stage acts on single pulse returns and, using Poisson statistics and 1st stage noise estimates, discards > 90% of the 1st stage residual noise.
- 3rd stage (algorithms completed, code under development) is designed to remove isolated noise counts in the immediate vicinity of actual target surfaces.





In Geiger Mode, the Avalanche PhotoDiode (APD) recovery time is very long (~50 nsec if actively quenched and ~1600 nsec if passively quenched). The probability of detecting a target beneath a canopy with one-way transmission T_c is given by

$$PD(n_s,\gamma) = \exp\left[-\frac{\gamma n_s}{2}\left(1-T_c^2\right)\right]\left[1-\exp\left(-T_c^2n_s\right)\right]$$

where n_s is the expected number of detected photoelectrons from the unobscured target,

 $\gamma = \frac{\rho_c}{\rho_t}$

and ρ_c and ρ_t are the reflectances of the canopy and target respectively. The second term in the equation gives the probability of detecting the signal with the Sigma multistop receiver while the first term gives the probability of disabling the receiver due to the detection of a canopy return. In contrast, the Sigma SPL detector /receiver has a very short recovery time on the order of 1.6 nsec and therefore only the second term in the equation is relevant.



The black curve in the plot shows the probability of detecting the unobscured target vs the mean signal strength. The red curve gives the same probability for the Sigma multistop SPL receiver in the presence of a 40% transmissive tree canopy ($T_c^2 = 0.16$). The remaining curves show the Geiger Mode probabilities for different values of γ . For $n_s = 5$, the probability of detecting the unobscured target is ~100%. The Sigma multistop receiver has a 57% probability while the single stop Geiger Mode receiver has only a 7% probability for $\gamma = 1$, even in the absence of solar noise (night operations) or detector dark counts.



SPL Surface Detection Probability vs Canopy Transmission and Unobscured Signal Strength

For 532 nm wavelength, Sigma SPLs are usually designed to generate $n_s = 3$ and 4 for green vegetative ($\rho_s = 0.1$) and dry vegetation/soil ($\rho_s = 0.15$) target surfaces respectively. Unlike Geiger Mode systems, more laser energy per beamlet always improves the surface detection probability.



Unobscured Signal Strength, ns











JIMO 3D Imaging Goals

•Globally map three Jovian moons

Horizontal Resolution: <5 m
Vertical Resolution: < 1 m

Worst Case Constraints

Europa (last stop) map must be completed within 30 days due to strong radiation field*
•348 orbits at 100 km altitude

•14.5 km mean spacing between JIMO ground tracks

•Surface Area: 31 million km²

* More recent JPL studies have indicated that, with proper shielding, Europa operations could possibly be extended to 3 or 4 months, allowing higher resolution maps.



Summary

•Our 100 beam scanning SPLs have provided decimeter level (horizontal) and few cm RMS (vertical) resolution topographic maps from aircraft AGLs up to 28 kft*. Data rates to date vary between 2.2 and 6 million 3D pixels per second., up to 60 times faster than conventional multiphoton lidars.

•The multibeam NASA MABEL pushbroom SPL has operated successfully at AGLs up to 65 kft

•Our smallest lidar, Mini-ATM, designed for cryospheric measurements, weighs only 28 pounds (12.7 kg), occupies 1 ft³ (0.028 m³), has a + 45 degree conical scan, fits in a mini-UAV, and covers more area with higher spatial resolution than the much larger and heavier predecessor NASA ATM system.

•Our low deadtime (1.6 nsec) detectors and range receivers permit daylight operation and multiple range measurements per pixel on a single pulse, allowing penetration of volumetric scatterers (tree canopies, water columns, ground fog, etc.).

•Our moderate to high altitude lidars built to date have been designed to provide contiguous topographic and bathymetric maps on a single overflight at aircraft speeds up to 220 knots (407 km/hr). Point densities and geolocation errors are compliant with USGS QL1 standards.

•We are currently flight testing the upgraded HRQLS-2 lidar which is designed to fly at altitudes between 11 and 15 kft. It will participate in USGS sponsored field trials later this month.

•Using a laser comparable to that developed for the SPL on NASA's ICESat-2 and a nominal 50 cm diameter telescope, one could globally map the three Jovian moons from a 100 km orbit with better than 5 m horizontal resolution in 1 month (Europa) or 2 months (Ganymede and Callisto) each. *Sigma customer for HAL has not yet given permission to show 28 kft data but spatial resolution is comparable to HRQLS-1 images at almost 4x the AGL. (28 kft vs 7.5 kft)



