# The 2009 Local ties survey at San Fernando Naval Observatory

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#### 1. Introduction

A local ties survey in the San Fernando Naval Observatory (ROA) was carried out during the summer of 2009. A SLR station (SFEL) contributing to ILRS and a Continuous GPS receiver (SFER), contributing to the IGS are collocated at ROA since 1996. Another CGPS receiver (ROAP) was installed a couple of years ago to contribute to the IGS Time Transfer Experiment. Goals:

to verify the old values

to complete the information linking not only these three reference points together but also linking them with other points to allow reviewing in the future. Disadvantages that comes from the situation of the main

points:

 the SLR station (SFEL) is located inside a closed dome at the top of the Observatory main building while the intermediate reference marks are placed at the main terrace (large height gradients).

to look for the IVP of the SLR telescope is not an easy task due to the reduced dimensions of the SLR telescope dome



Levels, EDM, Theodolites						
Instruments		Specifications				
NA2 Wild		σ = 0.3 mm/K	m. (double run)			
Leica TC2002 Total Station						
Leica TDA 5005 Total Station		$\sigma_D = 1mm+1ppm (DIN 18723)$ $\sigma_A = 0.15 mgon (DIN 18723)$				
Wild T3 SN134948		σ <sub>A</sub> = 0.2"				
GPS units.						
Station	Receiver		Antenna			
ROAP	SEPT POLARX2		SEN67157596+CR NONE			
SFER, RONE, RONW, TOAL	TRIMBLE NETRS		TRM29659.00 NONE			

Local Code	Local 8	Local & Global / IERS Designation			
SFR	SFER	13402M004	SAN FERNANDO IGS GNSS		
ROP	ROAP	13402M006	SAN FERNANDO IGS GNSS		
IVP (SFEL)	ROAP	13402M007	SAN FERNANDO SLR IVP		
PNE	RONE		SAN FERNANDO PILAR NE		
PNW	RONW		SAN FERNANDO PILAR NW		
TAL	TOAL		SAN FERNANDO TORREALTA		
Naming points table					

## 4. Ground Network and reference points representation

IVP (SFEL): the laser telescope is mounted on the top floor of the central tower that crown the main building of the observatory. The tower structure is independent of the rest of the building. The diameter of the circular plant where it is located is 5 m and it is covered and protected by a swivel me-tal hemispherical dome. The IVP is the intersection of the azimuth axis with the common perpendicular of azimuth and elevation axis of the San Fernando Satellite Laser Ranging telescope. The elevation axis comes from the knowledge of the height of the IVP from direct levelling observations.

The azimuth axis is materialized by two rotating targets (micro-prisms) at two different heights well above the SLR.

The determination of the elevation axis was carried out by the The determination of the elevation axis was carried out by the next method. They were placed two sheets of paper on both extreme circles, their centres were found (right and left) by different telescope elevation angles using two orthogonal azimuthally positions and finally the heights of these points here means the base of the second form these points and the second the second form the second term the second se were measured. The heights obtained from these centers have allowed us to establish that the axis is horizontal (HIVP). The resulting difference between them remains in the order of one to two tenths of a millimetre. Thus, the determination of IVP is the intersection of the azimuthal axis and the horizontal plane HIVP



SFR (SFER): type GNSS, the antenna is mounted on the end of an iron pole centered in the upper part of the so-called

SFER antenna has been intersected from the stations (PNW, PNE and TAL) and, in order to measure the distances from them Retro-tapes Leica 60x60 mm have been attached tangent to the tube. The point SFER (virtual) is still considered to 1.626 ROP (ROAP): type GNSS, the antenna is mounted on the top

of an iron pipe attached to the west facade of the "time department" building, and it stands out at the level of the terrace. The permanent mark is a eccentric bronze plate,

ROAP (antenna) and its eccentric mark on a bronze plate

(ROPm) cannot be occupied, therefore, they are obtained by measurements from other points of the network. **PNE (RONE), PNW (RONW), TAL (TOAL):** Concrete pillar protruding 1.5 m above ROA main building terrace and the mark consist of three brass centering devices.

ASF: Stainless steel plate with three brass centering devices. The mark refers to the centre of the three centering devices.

## 6. Data Analysis

The flow chart of the analysis processes is detailed in the figure on the left. First, classical and GPS raw data were processed and then a geodetic adjustment was made using GEOLAB. Second, the results were used as input of the AXIS 1.0.7 software to get the first azimuthal axis estimation. Third, a provisional IVP was calculated as intersection of this axis and the horizontal plane (HIVP), then a virtual elevation circle was estimated. Fourth, this new data was added and processed again using AXIS to get a quasi final IVP, and we go back to second step to get the IVP and axis values. Finally, rotation and translation of the 'geometrically modified' solution onto the required global reference frame was

## 7. Results

The last process provides the final azimuthal axis parameters and the local ties in ITRF2005. The results were sent to IERS for the new ITRF2008 in SINEX format. The results are summarize

be-	Dma	a target radius	0.2609 m	σ <sub>R</sub> =0.1	mm	low.
	Dmb target radius		0.0266 m	σ <sub>R</sub> =0.1 mm		
	Azimu	uth/Deflection	235º 26' 26.3"			
Loca	l Tie	ΔX (m)	ΔY (m)	ΔZ (m)	Dis	tance (m)
0	LD	-45.0950	35.2995	89.5475	1	06.2938
NE	w	-45.1017	35.2780	89.5560	1	06.2966
Prec	ision	0.0017	0.0009	0.0020		0.0027
Diffe	rence	-0.0067	0.0215	-0.0085		

References: http://www.iers.org/nn\_10900/IERS/EN/Organization/Working Acknowledgments: We thank to Jon Dawson to provide us the last Axis software release, to 'ETSI en Toografia, Geodesia y Cardografia de Madrid' for lent us devices and to F. de la Cruz, J. Menéndez and L. Espiga for their hard work.

### 2. Site Description

San Fernando Observatory is located in the South of Spain near Strait of Gibraltar. The SLR is co-located with two GNSS observation system stations. There are a number of survey monuments and pillars within the observatory which serve as reference marks for the local tie determination. It is from this network of survey marks and pillars that terrestrial measurements have been made to obtain the 3D coordinates, firstly in a local system, later they were transformed to the Global system.



### 5. Observations

The "local network" consists of the so-called "external network" and the "internal network". The second one, for the determination of the station SFEL IVP (ILRS), comprising sequences  $\,$  DMan and DMbn (n = 1, 2, ...,8) and the altimetric points in the dome



was conducted among the survey points, using on one hand the elevation angle and slope distances with EDM and on the other hand height differences with Leica precision levels and invar rod. Levelling loops covering all monuments were completed in both directions. SLR

Two posts fixed to the telescope were deployed on their tops. Coordi-nates were obtained at eight azimuthal positions of the telescope measuring directions and distances from PNW, PNE and TAL. Thus the radius of the circle described by the lowest prism DMb was ~ 26 mm and the uppest prism DMa was ~ 266 mm. GPS

GPS survey has been made in order to give results within ITRF geocentric frame and also to align the local system. PNE, PNO, TAL and of course ROAP and SFER IGS stations were observed.

Raw Terrestrial Observations		Raw GPS Observations
Measurement correction. Calibration. Atmosphere. Collimation.		GPS calculus & analysis BERNESE
Height of instrument Determination.		
	,	Ļ
Coordinates	Adjustment and their variance-covariance GEOLAB	
	,	
Extraction of Geolab results from bin the va	ary output including geodetic p riance-covariance matrix. DMPGEOLAB	parameter estimates and
	,	
System geometrical modeling. Azim	uthal axis parameter estimatio Residual analysis. AXIS	n. Solution adjustment.
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Determination of provisio	nal IVP and virtual elevation cir	rcle (4 points)
		Ť
Adju Coordinates and the GE	stment. ir variance-covariance. OLAB	
Extraction of Geolab results from bin estimates and the var	ary output including geodetic p iance-covariance matrix.	arameter
Dirit.		
System geometrical modeling. IVP e axis and virtual elevation axis. S Residual analys AXIS	estimation with azimutal olution adjustement.	Final
		YES
Transformation of coordinates varia read	nce-covariance into Earth cent ness for SINEX output. AXIS	red Cartesian system in
Alignment of arbitrary terrestrial ne	twork to the International Terr AXIS	estrial Reference Frame via transforr

a priori AXIS

## 3. Instrumentation