ILRS QCB Meeting

January 19, 2021 Agenda

- Brief on the ILRS contribution to the Reference Frame (Erricos)
- Minimum FR population for NPs on LEO satellites using LARES (John)
- Issues regarding polynomial fits and clipping (Matt)
- The Simosato Story (Van)
- Any updates in Minico results (Van)
- Do any of the satellite C/M models need to be changed to accommodate the new station configurations (Jose)
- More stress on long and short stability rather than NP How do we implement this?
- History Change Logs
- Document on Best practices (calibrations, barometer, etc.). (Matt)
- Anything else?

Low-return Normal Point Analysis

John C. Ries 1/19/2021





The University of Texas at Austin Cockrell School of Engineering

- Daytime normal points minimum 6 data points
- Night time normal points minimum 3 data points
- Fewer data points would be acceptable on lower satellites (5-second normal points) from those ranging systems with lower pulse repetition rates where these minimum requirements are not practical.
- What is the impact of making NPTS with as few as 1 return?
 - In the following, 'low-return NPTs' refers to NPTs with less than 3 returns, though the analysis was extended to 'lowerreturn' NPTs with less than 6 returns.



Breakdown of NPTS by number of returns

(January 2020 for LAGEOS)

-								
	STATION	1SH0T	2SH0TS	3SH0TS	4SH0TS	5SH0TS	6+SH0TS	
	1873	4	5	5	4	2	28	A small number of
	1884	0	0	0	0	1	27	stations are responsible
	1888	0	0	0	0	0	45	for the low- and low-
	1890	0	0	0	0	0	129	return normal points.
	1893	5	6	4	5	4	47	
	7090	64	62	44	44	42	703	Look closer at
	7105	3	4	4	6	6	268	Yarragadee to test impact
	7110	15	8	5	10	4	282	of successively removing
	7119	7	7	4	3	3	101	NPTS with only 1 shot,
	7237	0	0	0	0	0	301	then 2 shots,up to 5
	7249	0	0	0	0	0	17	shots, since it has the
	7501	4	3	1	1	0	94	most low- and lower-
	7810	0	0	0	1	5	1036	
	7811	0	0	0	0	1	97	return NPTs, and thus
	7821	0	0	0	0	0	90	should be impacted the
	7825	0	0	0	0	0	44	most.
	7827	0	0	0	0	0	218	
	7838	3	2	5	7	8	216	
	7839	0	0	0	0	0	307	
	7840	0	0	0	0	0	478	
	7841	0	0	0	0	0	144	
	7845	0	0	0	0	0	392	
	7941	7	4	6	2	2	479	
	8834	0	0	0	0	0	230	





Fit and NPT precision statistics (in mm) (only a small number of orbit parameters are estimated)

CASE	TOTAL OBS	FIT RMS	B/TB RMS	POLY RMS	test1 contains all NPTS (from all stations)
test1 (7090	only) 959	7.6	2.9	2.4	but the results
test2	× 895	7.4	2.4	2.0	for 7090 are
test3	833	7.2	2.4	1.8	shown
test4	× 789	7.2	2.4	1.8	SHOWIT
test5	745	7.2	2.3	1.6	1 10
test6 Y	× 703	6.5	1.9	1.6	test2 uses NPTs with at

Moving from bottom to top, as NPTs with fewer and fewer returns are included, the FIT RMS degrades. The POLY RMS (the estimated NPT precision) also degrades, indicating that the scatter is significantly worse for NPTs with only 1 or 2 returns. NPTs with at

A few passes are lost if low-return NPTs are excluded, but these are going to be the most unreliable in any case.

However, the impact on the position estimates for 7090 was minor; no difference larger than 0.6 mm was observed for any component (ENU) for any case (31-day estimate; LAGEOS-1; Jan 2020).

Orbit differences were generally not very significant; low-return NPTs are not a large percentage of the overall tracking. RMS orbit differences were 2-3 mm (RMS) for L1/L2 and 5-7 mm (RMS) for LARES and Starlette.





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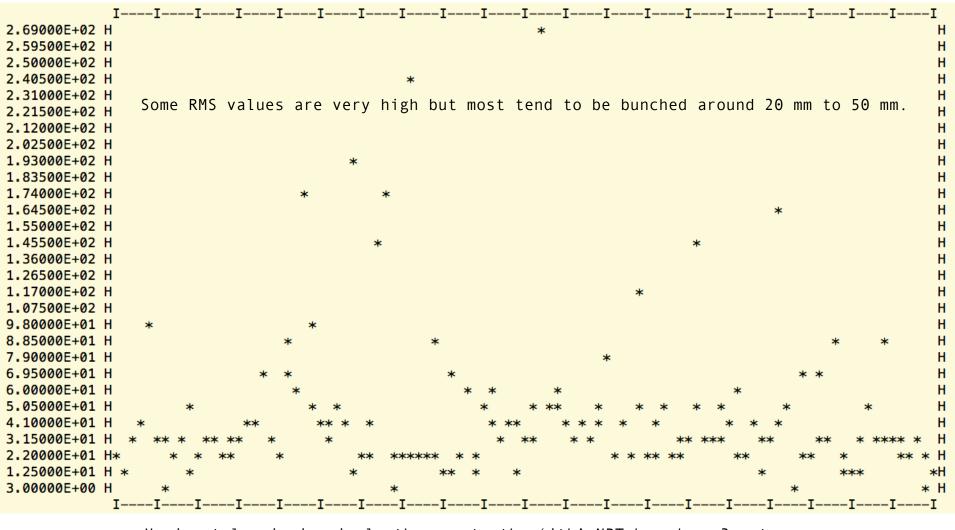
least 3 returns

Similarly up to test6, which includes only

NPTs with at

least 6 returns

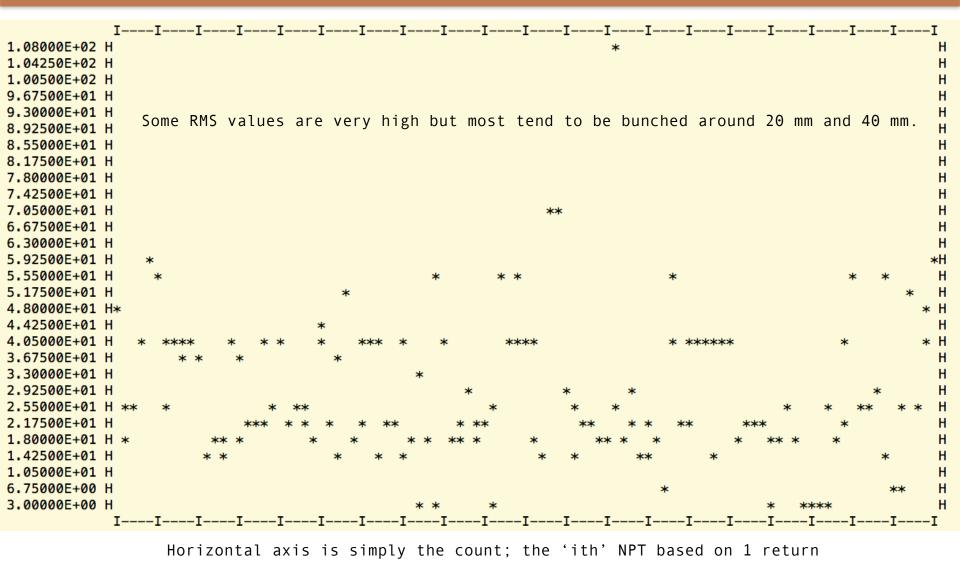
Assigned NPT RMS for 2-return NPTs



Horizontal axis is simply the count; the 'ith' NPT based on 2 returns



Assigned NPT RMS for single-return NPTs





Fit statistics for whole network

SLR data for all of 2020 fit using all available NPTs for 4 satellites at various altitudes. Residual RMS was computed for 7 cases: The reference fit RMS used all NPTs, then the RMS was computed for just the 6+ return NPTs, just the 5-return NPTs, just the 4-return NPTs, etc. down to the single-return NPTs.

RMS	All NPTs		6+ Returns		5 Returns		4 Returns		3 Returns		2 Returns		1 Return	
	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)
LAGEOS-1	61550	7.1	57708	6.9	634	8.7	645	8.8	727	8.7	834	9.6	1010	10.1
LAGEOS-2	55242	7.4	51002	7.2	738	8.4	711	8.8	811	9.4	895	9.1	1066	10.1
Starlette	83009	27.4	77999	27.0	693	29.2	797	28.3	909	28.5	1050	31.0	1358	31.1
LARES	70071	19.3	63811	18.5	839	21.8	987	21.7	1115	23.5	1369	24.6	1796	24.2
7105														
LAGEOS-1	2913	7.7	2656	7.2	42	9.5	34	10.6	43	12.6	65	11.0	64	11.8
LAGEOS-2	2786	7.1	2452	6.3	49	8.8	48	10.2	62	9.8	71	10.5	92	10.9
7090														
LAGEOS-1	7693	7.2	5870	6.8	275	7.9	316	7.8	333	8.0	388	8.3	506	9.1
LAGEOS-2	6999	7.2	5006	6.9	347	7.2	338	12.6	361	7.9	434	8.0	504	8.5
Range Bias	All NPTs	6+ Returns	5 Returns	4 Returns	3 Returns	2 Returns	1 Return							
LAGEOS-1	0.7	0.7	0.9	1.5	1.2	0.8	0.6	Range bias = network weighted average of residuals (mm) for January-November 2020						
LAGEOS-2	1.7	1.6	2.9	2.4	2.5	2.2	2.1							
Starlette	0.0	0.0	10.0	7.9	5.0	6.9	6.2	(SLRF2014 and ILRS 2013 station-dependent CoM used)						
LARES	-0.8	-0.4	-6.6	-4.1	-5.7	-3.9	-5.7	(7-day arcs for L1/L2/LARES, 6-day arcs for Starlette)						

There is a significant increase in the noise for all targets, with 1- and 2-return NPTs not surprisingly the worst. Effect was especially apparent for 7105.

For LARES and Starlette, the lower-return NPTs appear to be significantly biased compared to the 6+ return NPTs. No apparent issue for L1/L2.



Downweighting Experiment

Using a simple rule, NPTS were downweighted if they had less than 6 returns. This retains all the data, including passes that are dominantly low-return normal points (small differences in the number of NPTs is due to slightly different points being edited).

RMS	All NPTs		6+ Returns		Downwt	
	# NPTs	RMS (mm)	# NPTs	RMS (mm)	# NPTs	RMS (mm)
LAGEOS-1	68557	7.3	63969	7.1	68558	7.2
LAGEOS-2	60246	7.4	55342	7.2	60237	7.3
Starlette	92042	27.4	86143	27.0	91977	26.7
LARES	75590	19.4	68170	18.6	75431	18.3
7105						
LAGEOS-1	3358	7.7	3077	7.3	3359	7.7
LAGEOS-2	3148	7.2	277	6.5	3147	7.3
7090						
LAGEOS-1	8713	7.3	6616	6.9	8713	7.5
LAGEOS-2	7966	7.1	5693	6.8	7966	7.3

While this simple rule provides some small improvement in the overall RMS for all satellites, it does not necessarily benefit every station. A more sophisticated rule might perform better.



Conclusions (1)

- The FIT RMS increases when low-return NPTs are included.
 - The low-return NPTs are clearly worse than NPTs with at least 6 returns, even in the case of 5, 4 or even 3-return NPTs.
 - For the two smaller (and lower) satellites looked at, a significant bias of 5-10 mm is introduced for NPTs with less than 6 returns.
- There seems to be some inconsistency in computing the RMS for low-return NPTs.
 - Since the uncertainty of a low-return NPT is large, the assigned RMS should probably be correspondingly large.
 - However, the analysts do not generally use it to inform their data weighting in any case, so there does not seem to be much point in trying to impose strict requirements on how to assign the RMS.





Conclusions (2)

- The geodetic impact of the low-return NPTs appears small, but analysts should consider the impact of including NPTs with less than 6 returns.
 - Analysts may want to test the impact of including/excluding any NPT with less than 6 returns, particularly for low satellites, to see the effect on the biases (e.g., Jason-3).
 - Some passes would be lost, but these are also clearly among the least reliable.
 - Where the tracking coverage is already weak, keeping the less reliable NPTs may be better. Keeping them but downweighting them may be reasonable.
 - This should become less of a problem with time as high-rep-rate stations come on line (7840, for example, had only a single NPT with less than 6 returns even when the test NPT software was run to allow for NPTs from as few as a single return).
- A possible analysis strategy would be to include NPTs with less than 6 returns, but downweight them progressively more severly as the number of returns gets smaller.
 - Preserves passes or low-elevation data that might otherwise be lost
 - Testing a simple down-weighting scheme gave mixed results





Range Residuals to Normal Points - How flat is too flat?

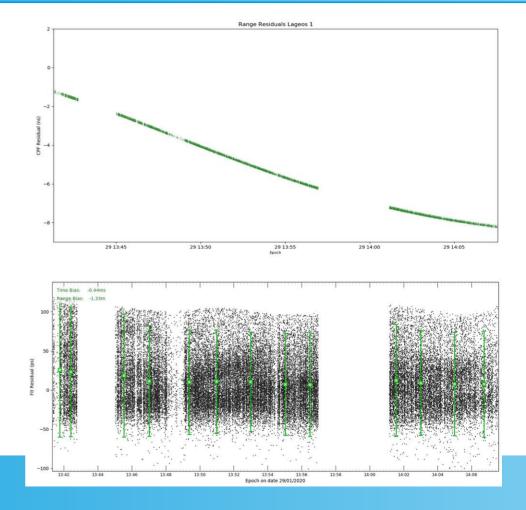
Matthew Wilkinson SGF, Herstmonceux

Extracting SLR Returns

SLR ranges are plotted as residuals to a reference orbit.

These O-C residuals must then be flattened in order to form Normal Points.

This is achieved at stations using either orbital adjustment or a high order polynomial fit.

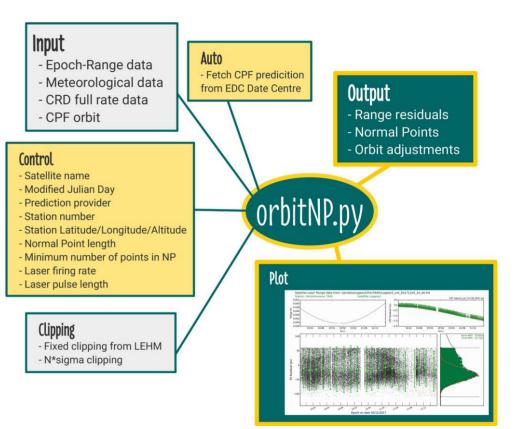


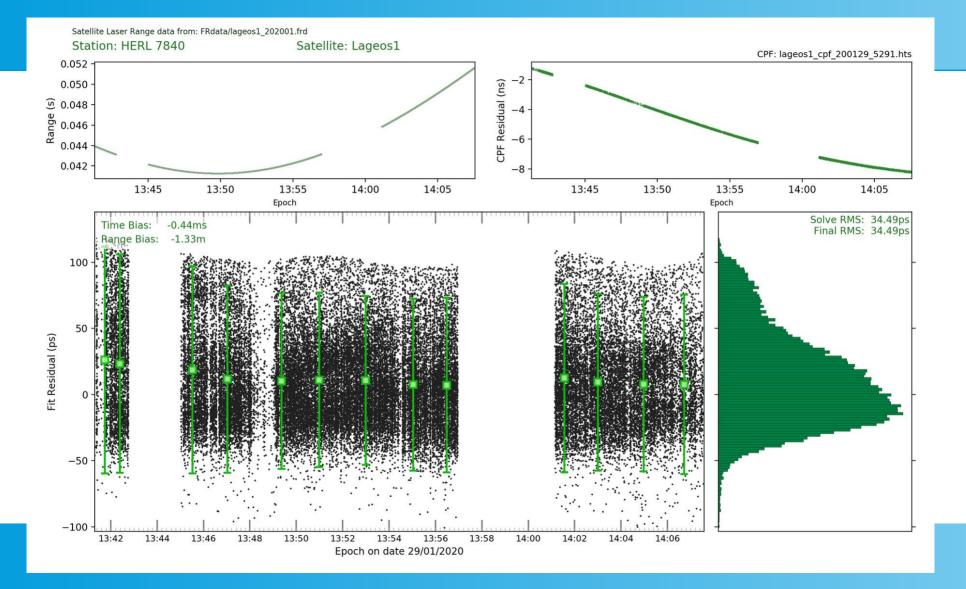
OrbitNP.py

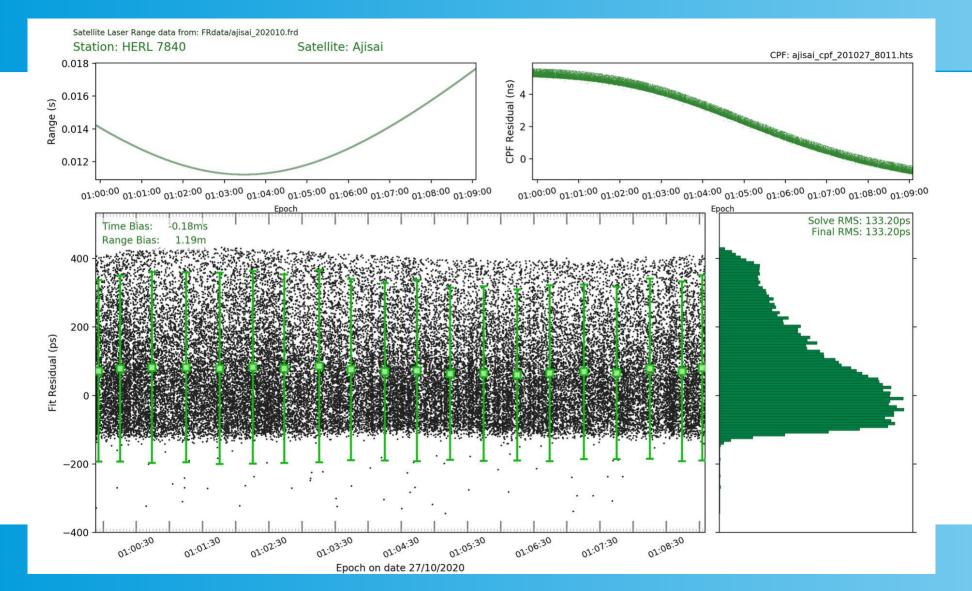
Flattening by orbit adjustment can be achieved using the **orbitNP.py** software.

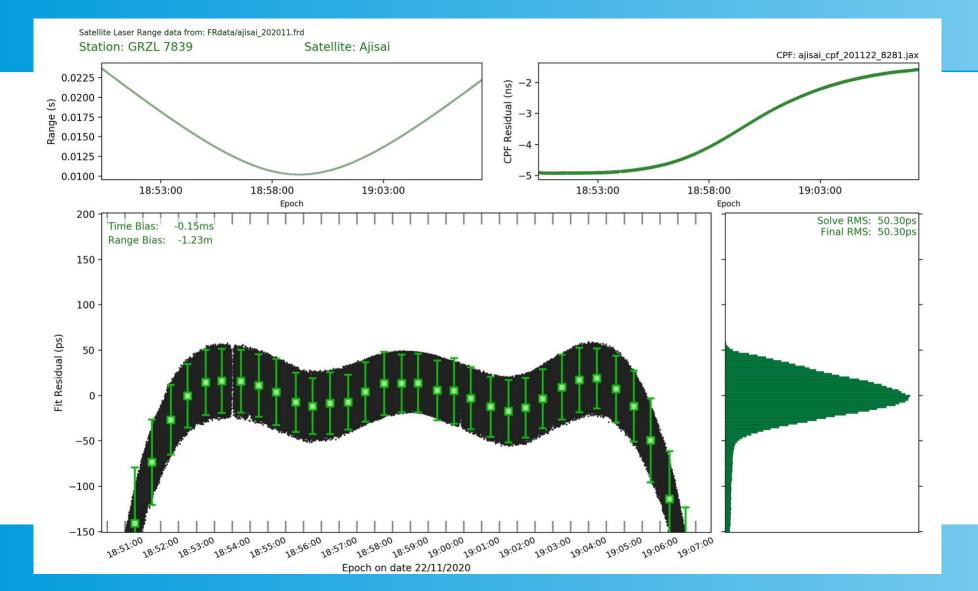
Using this software, it is possible to pick out and process individual passes from full-rate data for all stations and satellites.

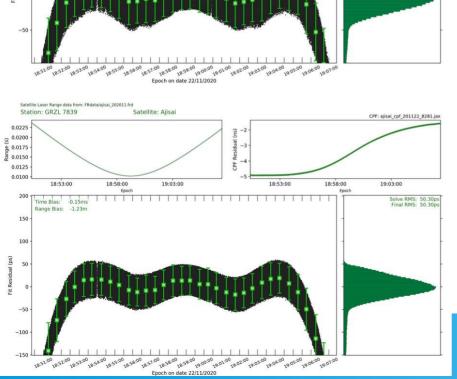
Each station decides the method to form residuals and the data clipping criteria.

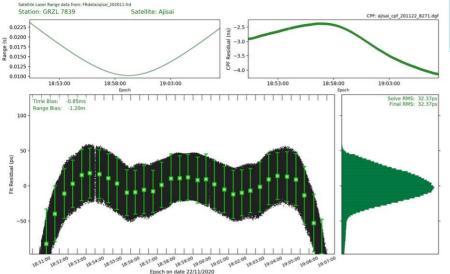






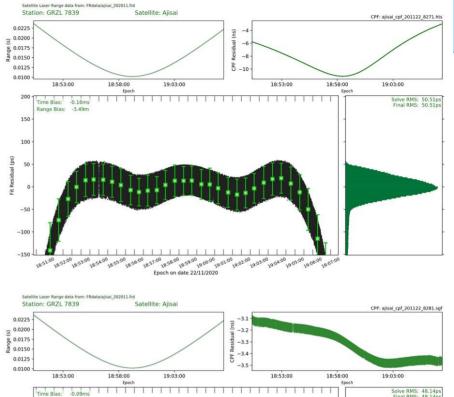


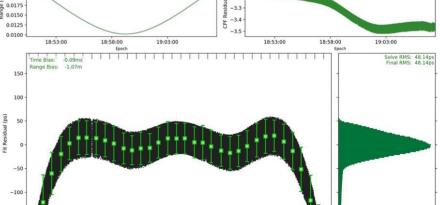




Ajisai

- Graz



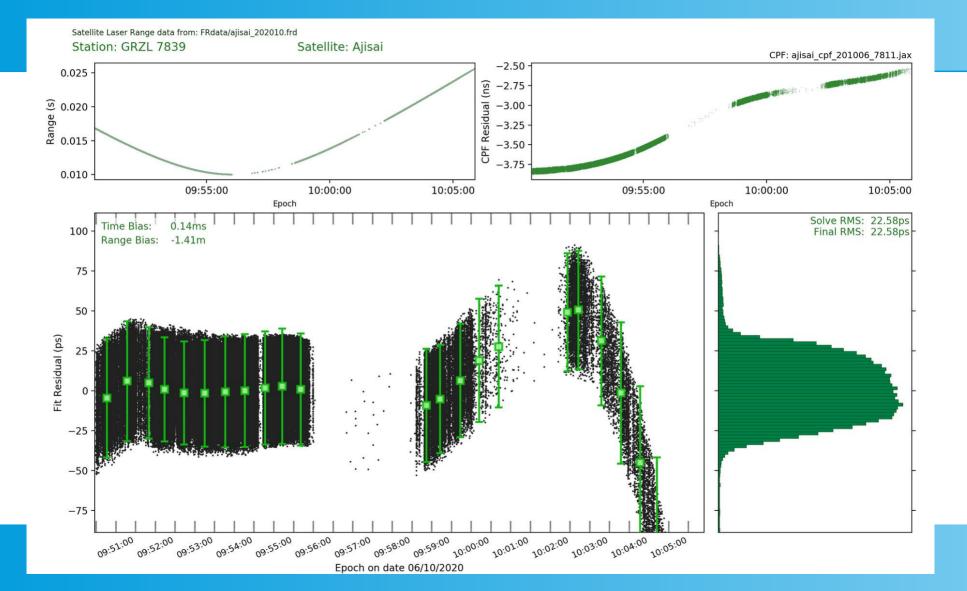


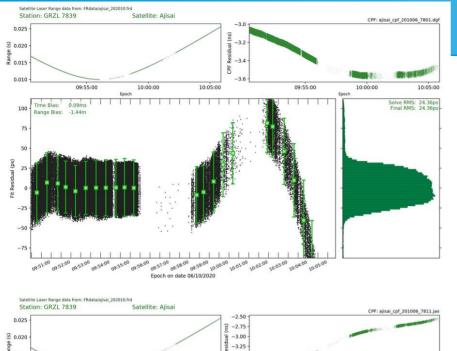
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28:59:00

Epoch on date 22/11/2020

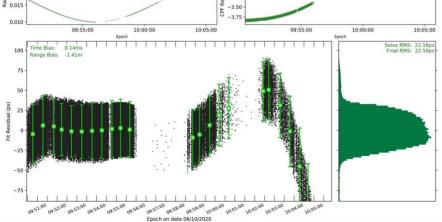
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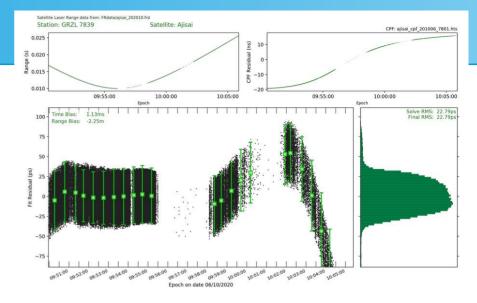


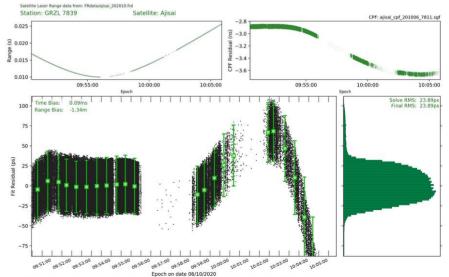


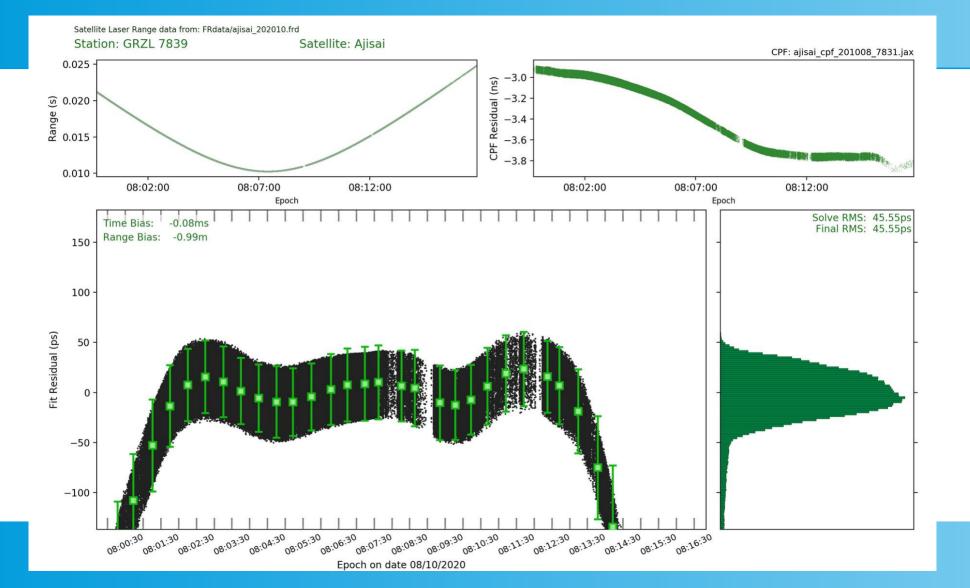
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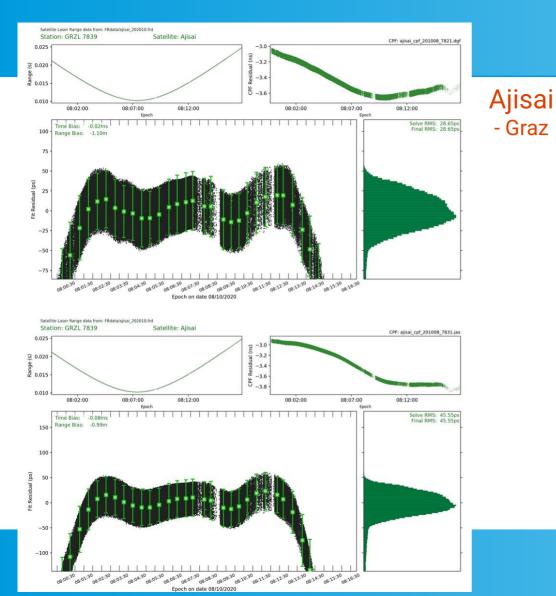
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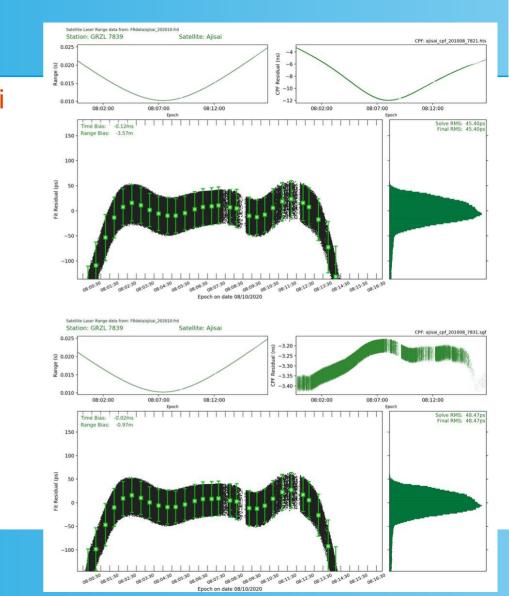


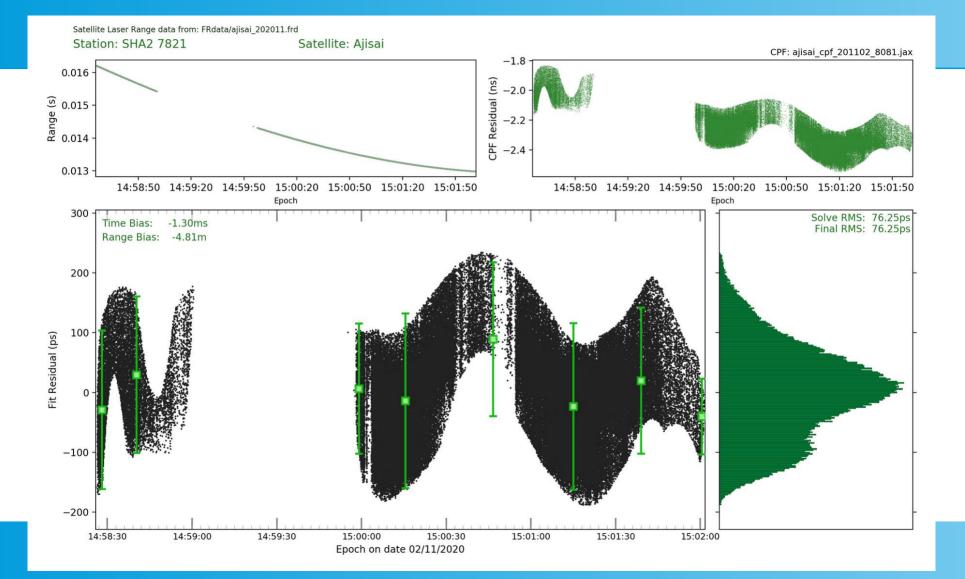


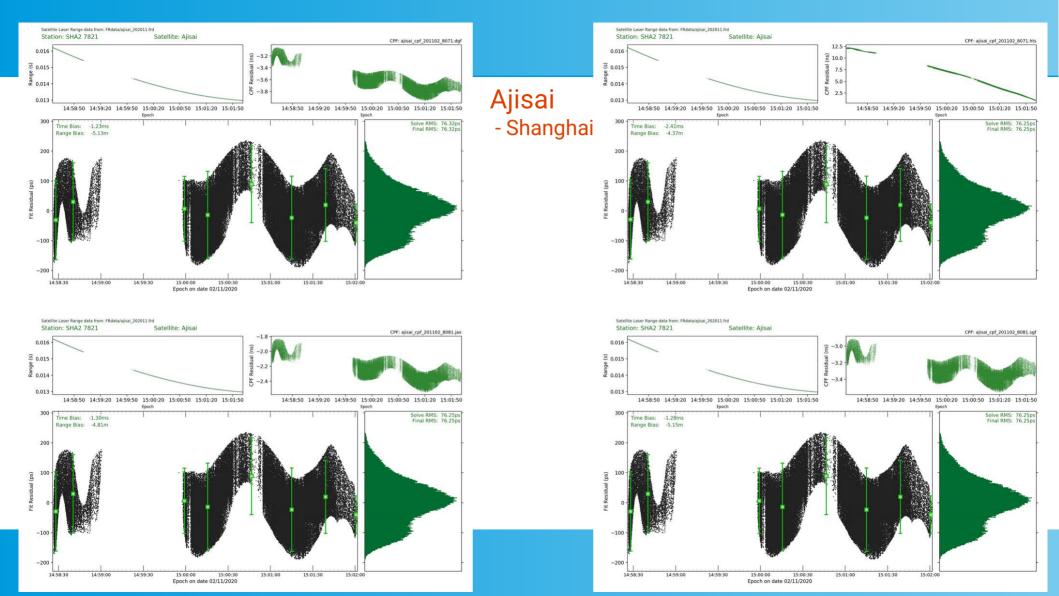


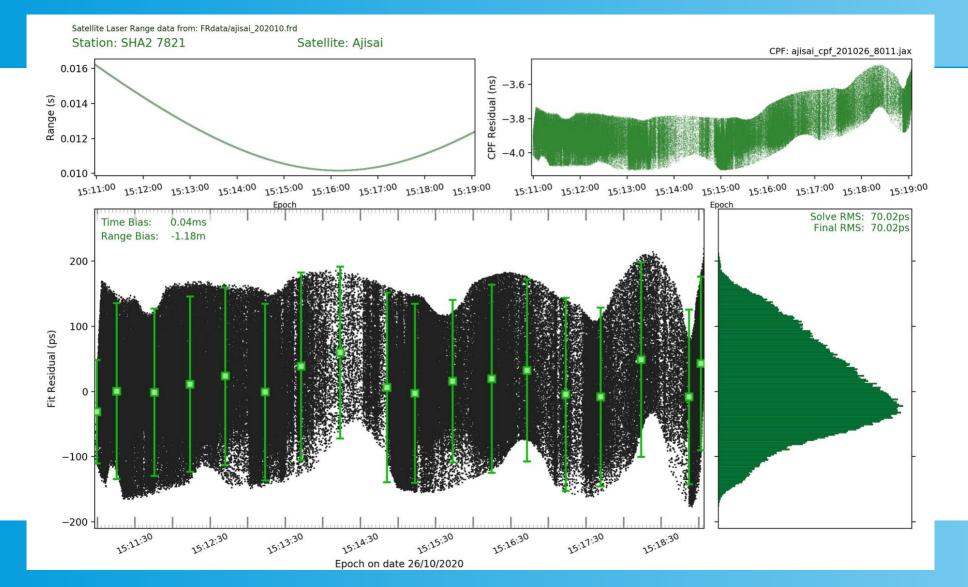


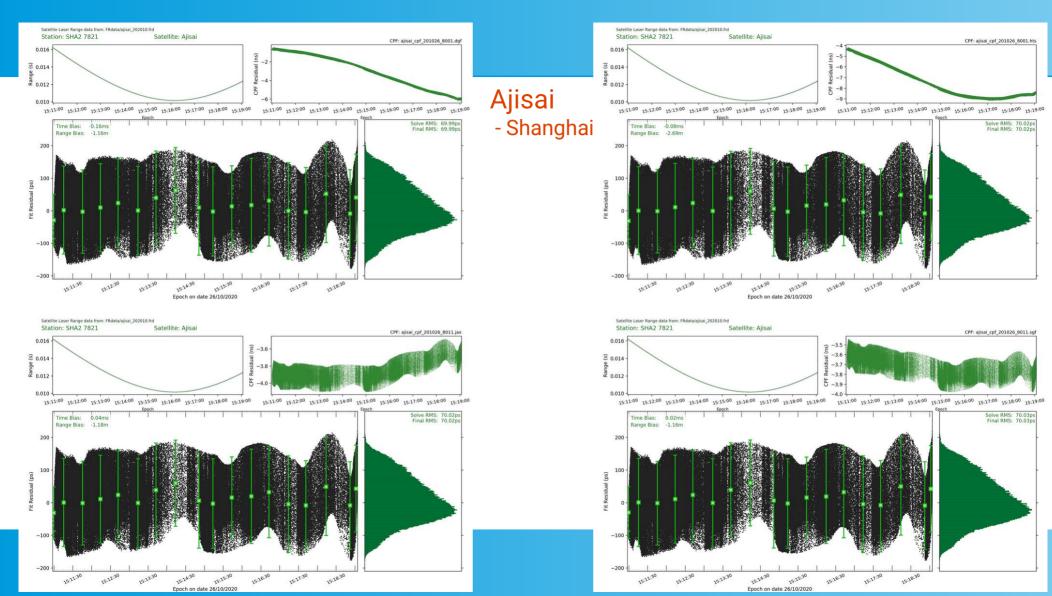


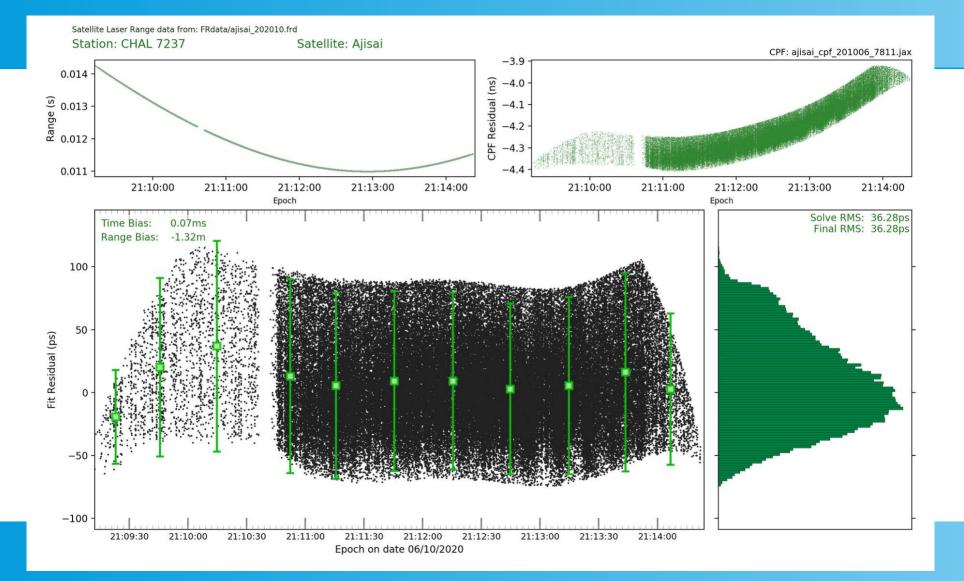


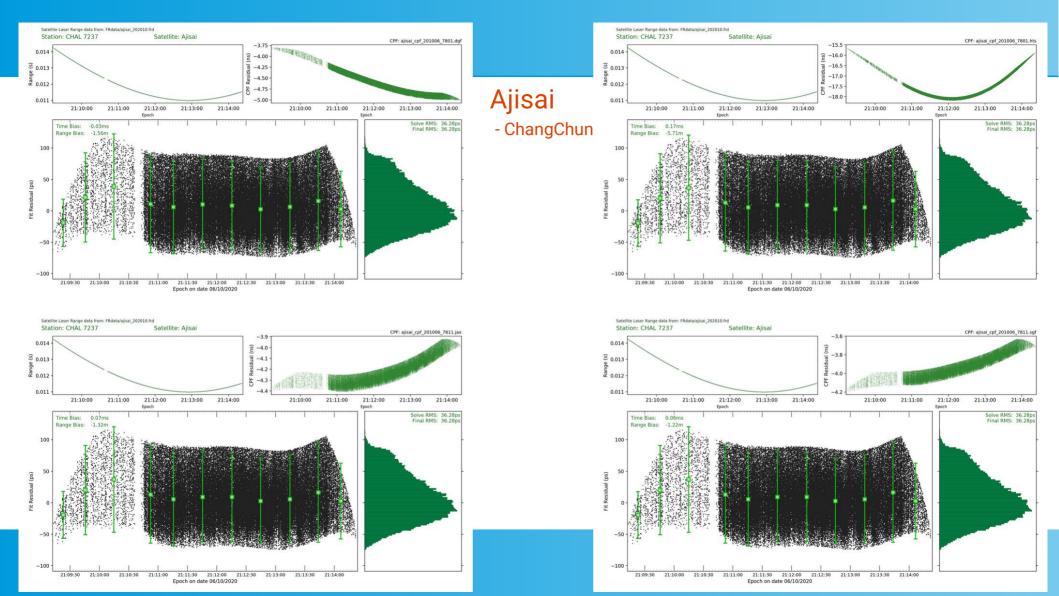


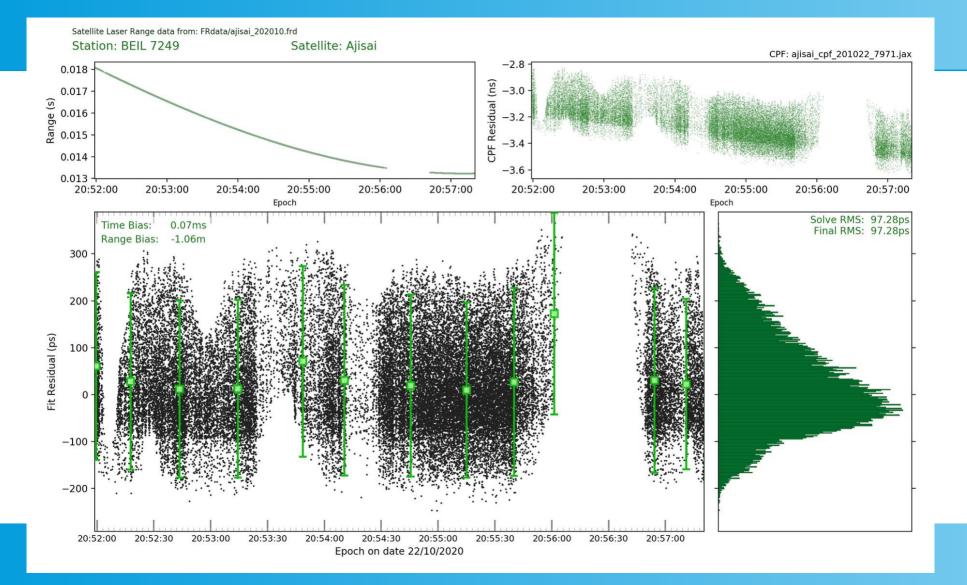


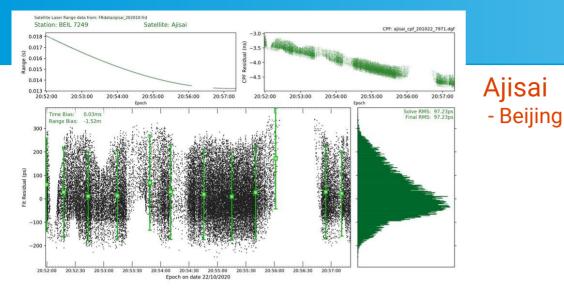


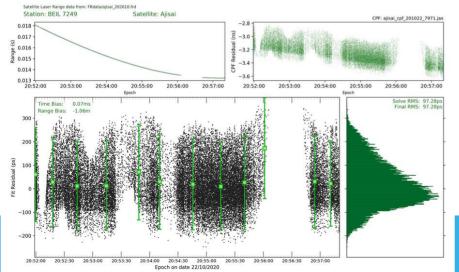


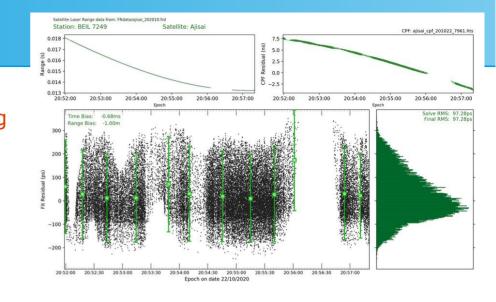


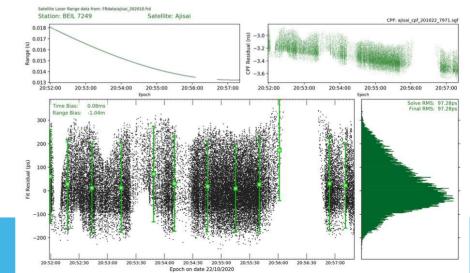


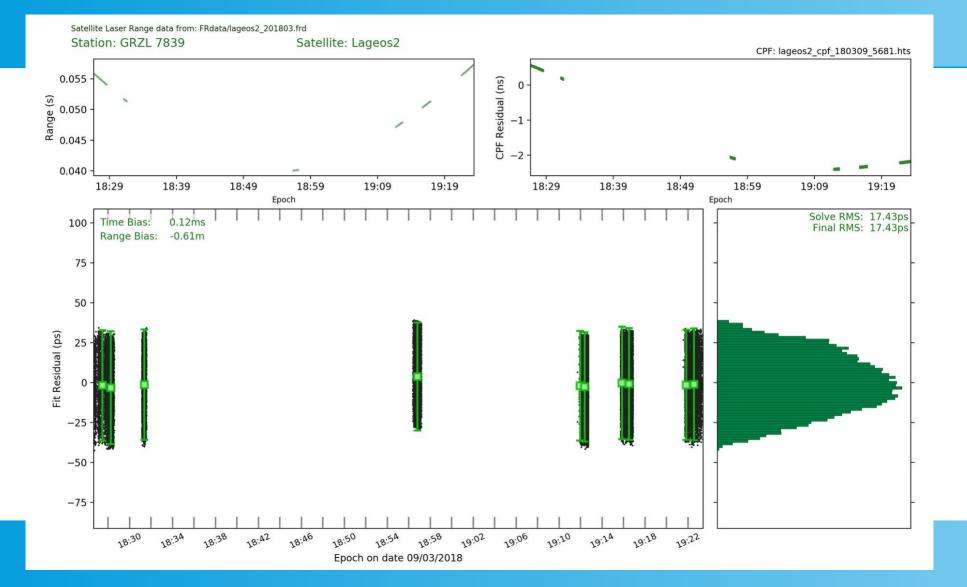


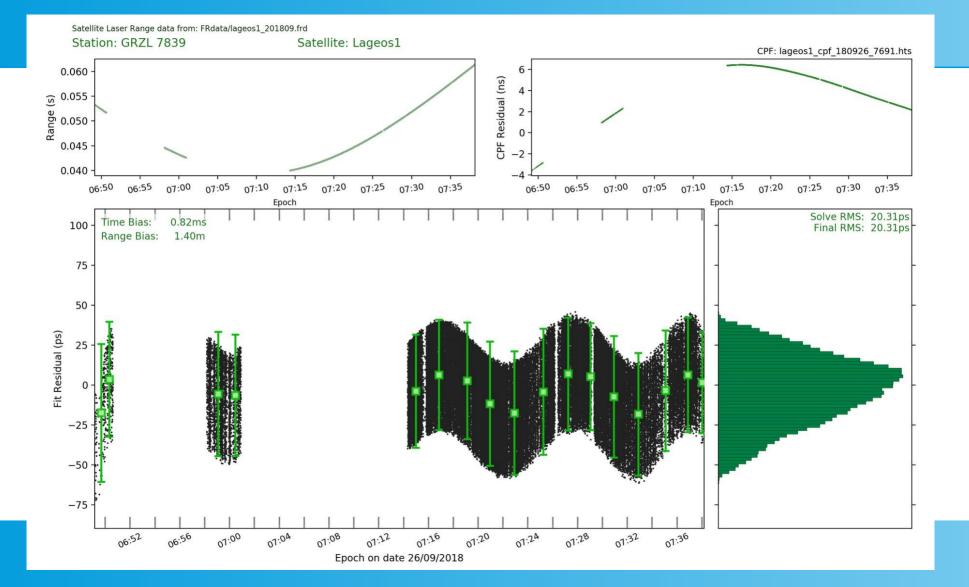


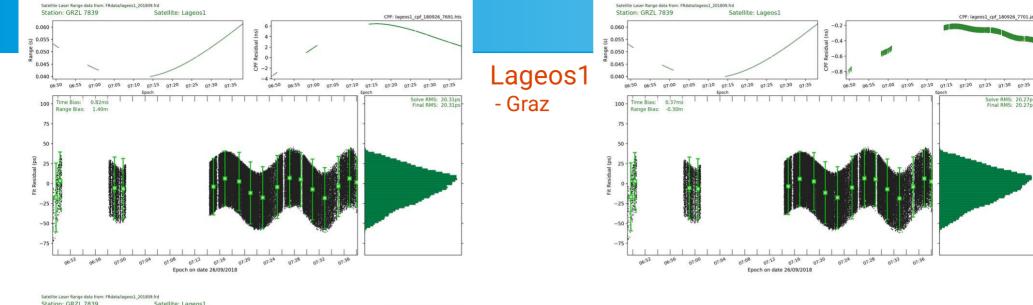






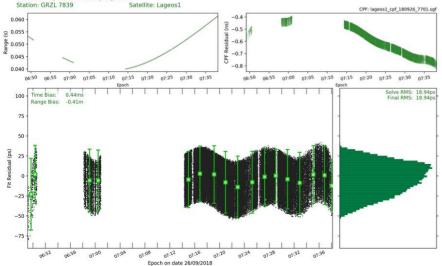


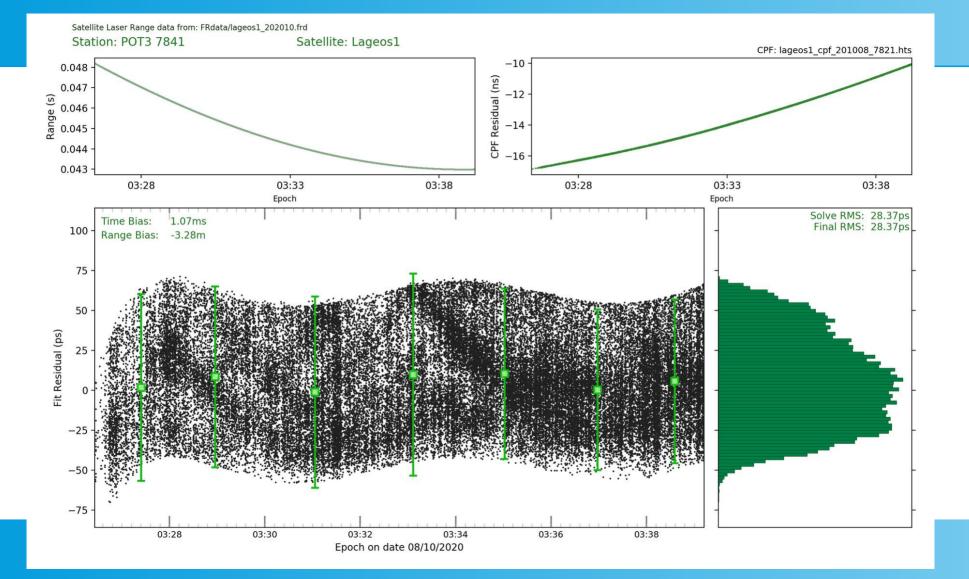


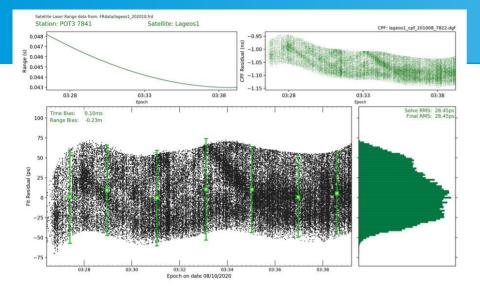


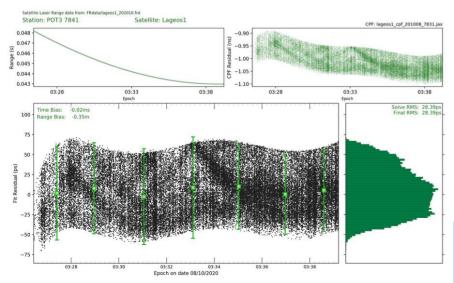
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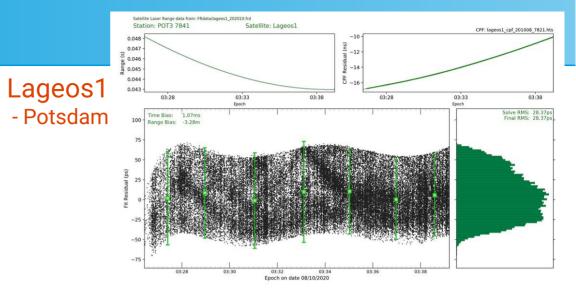
Solve RMS: 20.27ps Final RMS: 20.27ps

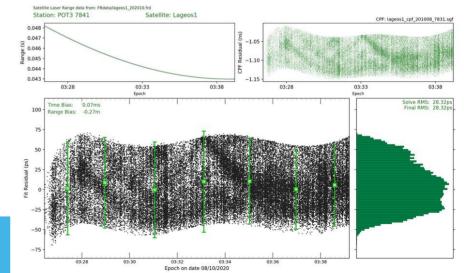


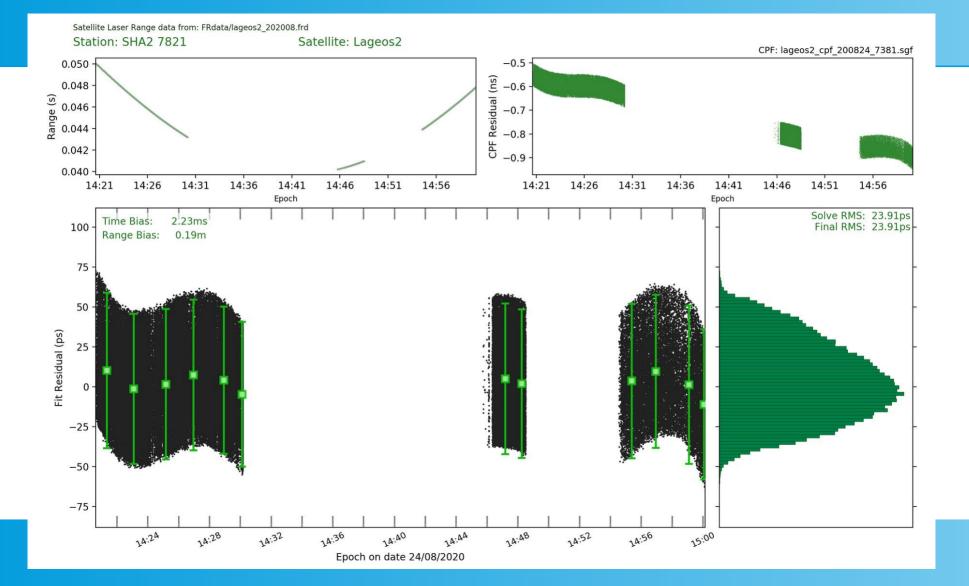


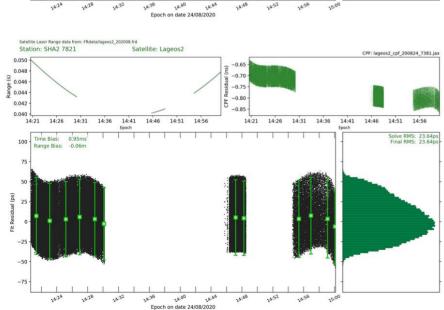


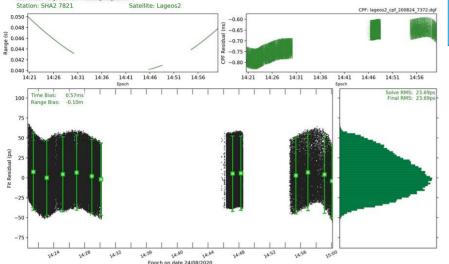




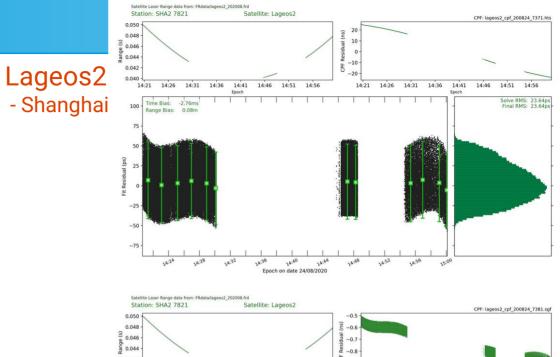


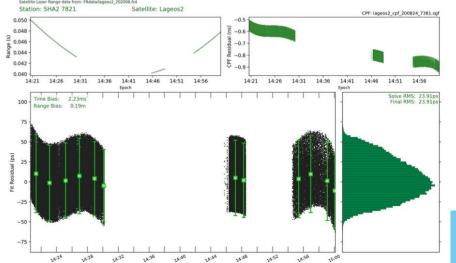






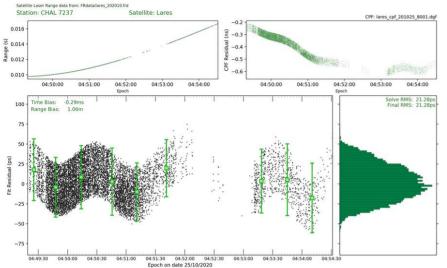
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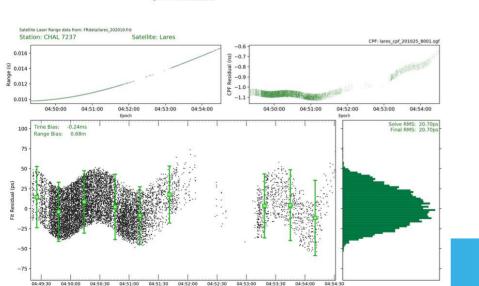


Epoch on date 24/08/2020

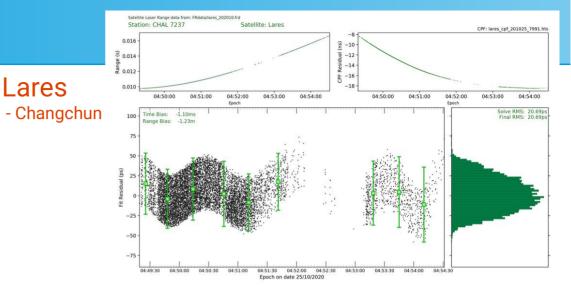
Satellite Laser Range data from: FRdata/lares_202010.frd Satellite: Lares Station: CHAL 7237 CPF: lares_cpf_201025_7991.hts -8 0.016 CPF Residual (ns) -10(s) 80.014 (s) 80.014 -12 -14 -160.010 -1804:50:00 04:51:00 04:52:00 04:53:00 04:54:00 04:50:00 04:51:00 04:52:00 04:53:00 04:54:00 Epoch Epoch Solve RMS: 20.69ps Final RMS: 20.69ps Time Bias: -1.10ms 100 Range Bias: -1.23m 75 50 Fit Residual (ps) 25 C -25 -50 -75 04:49:30 04:50:00 04:50:30 04:51:00 04:51:30 04:52:00 04:52:30 04:53:00 04:53:30 04:54:00 04:54:30 Epoch on date 25/10/2020

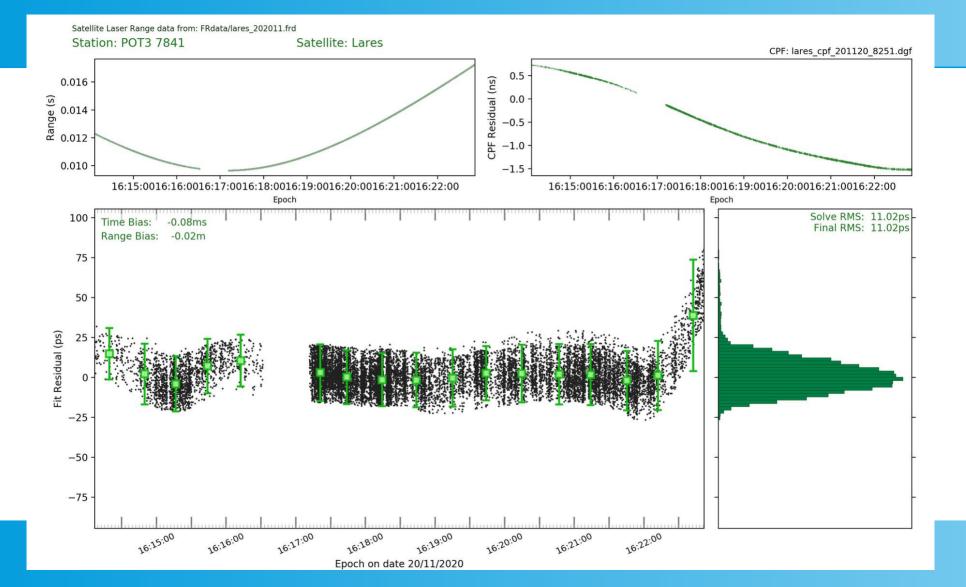


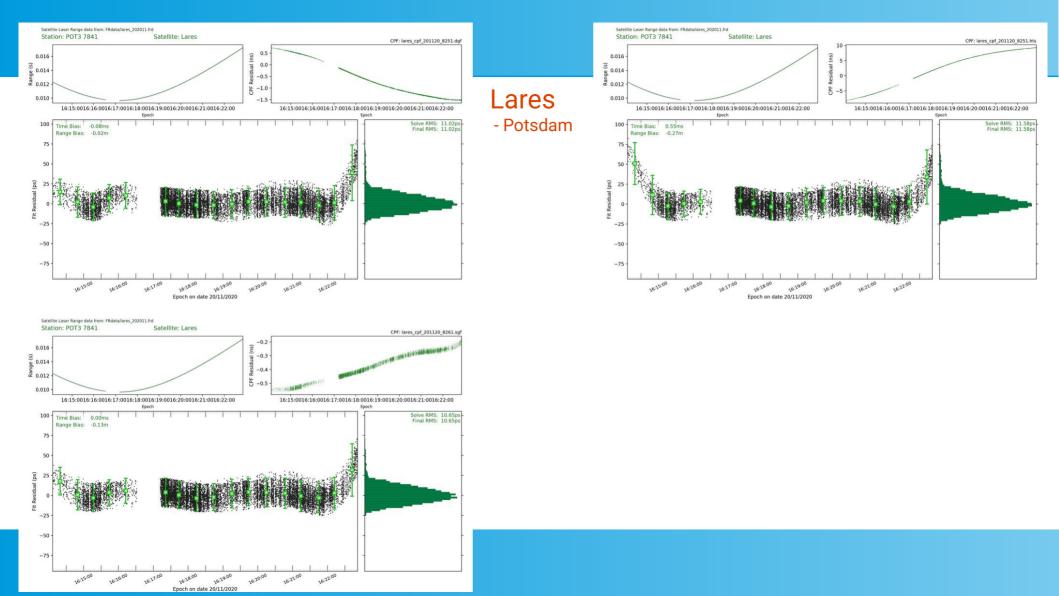
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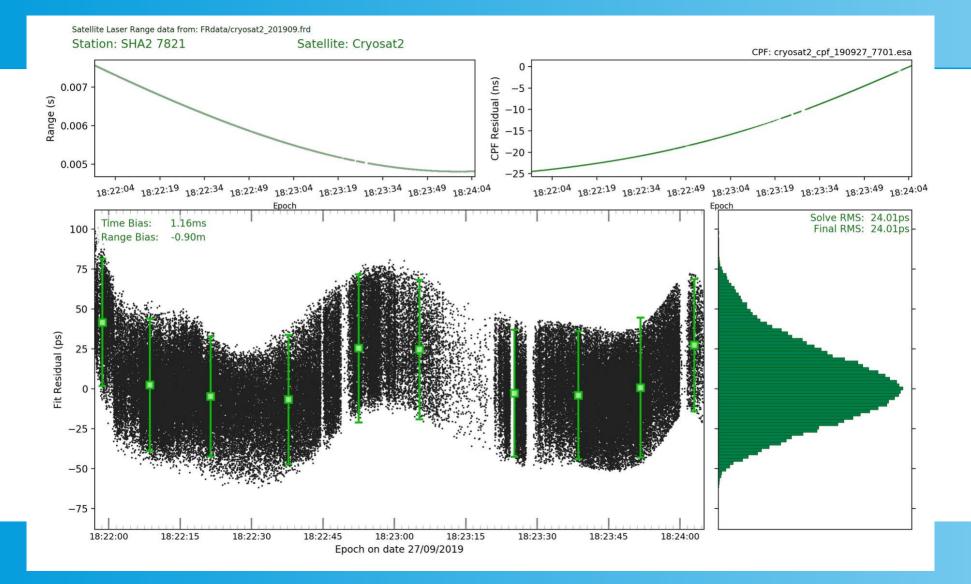


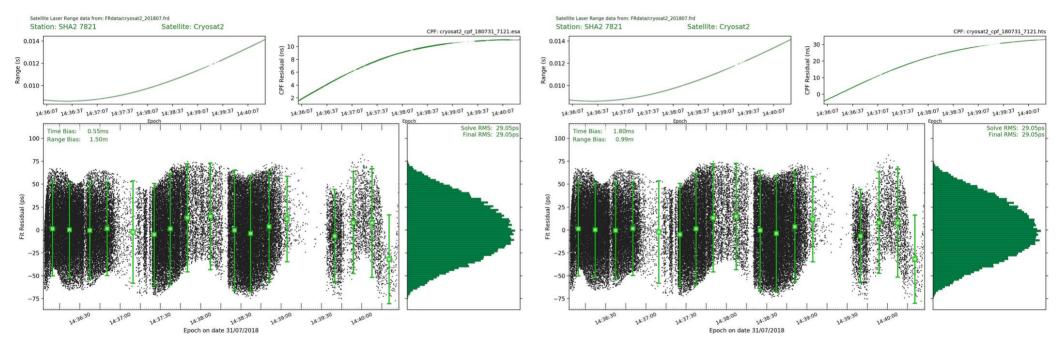
Epoch on date 25/10/2020

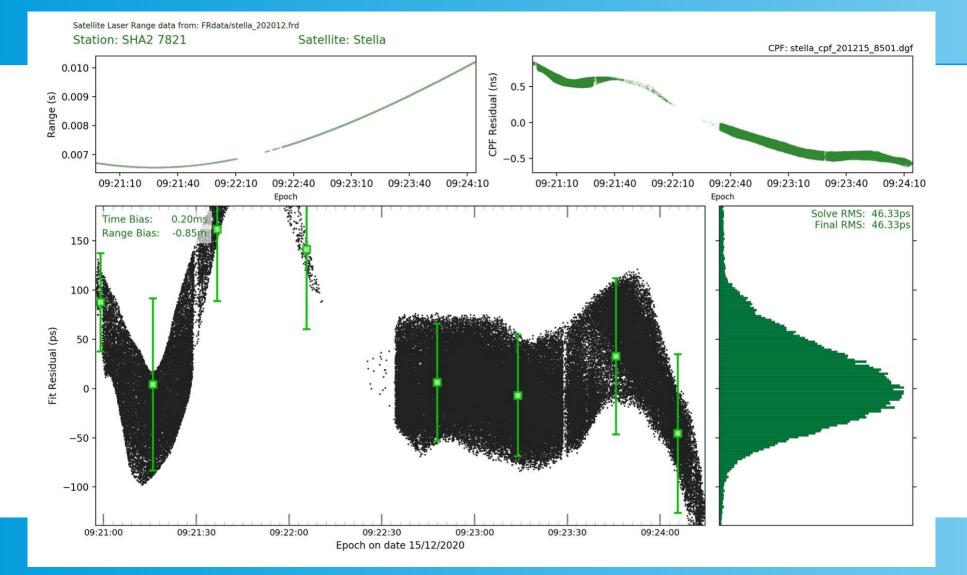


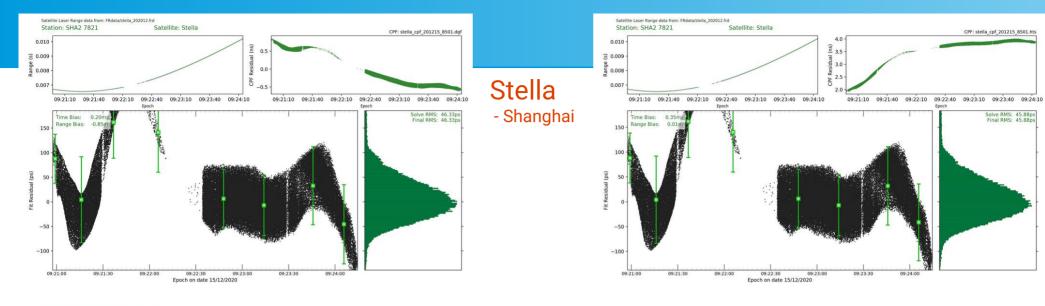


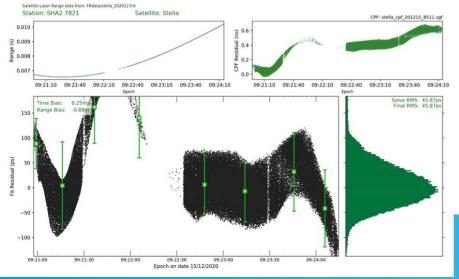




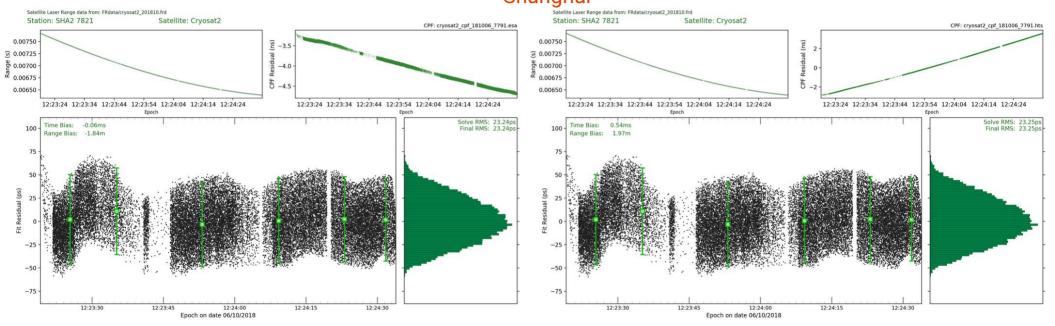


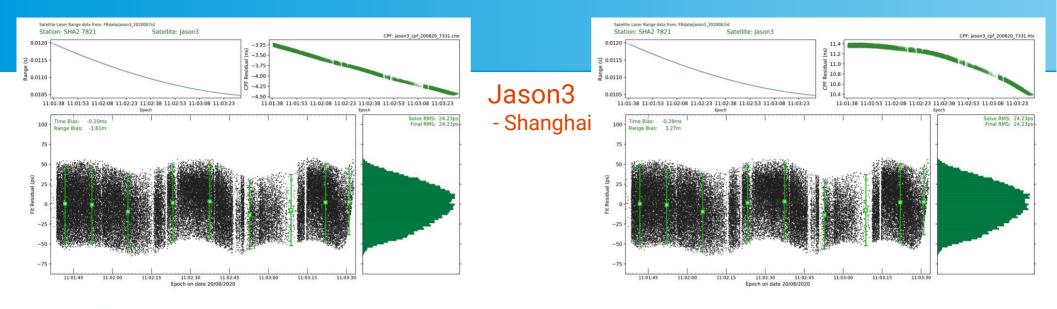


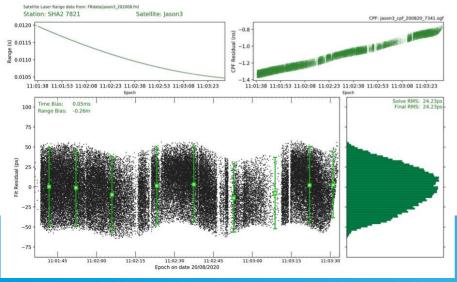


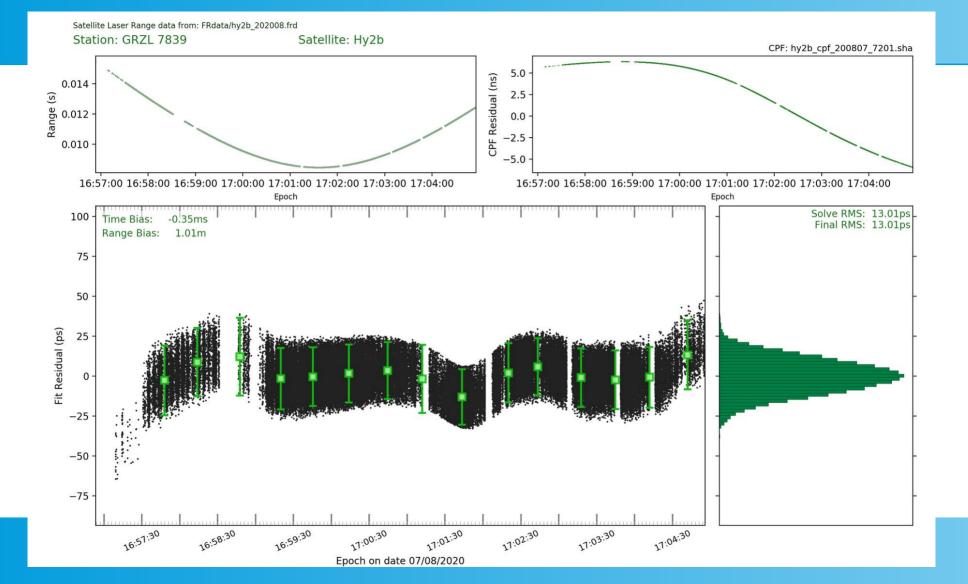


Cryosat2 - Shanghai settle targe data from: Fridata(ryosat2,201810,frd

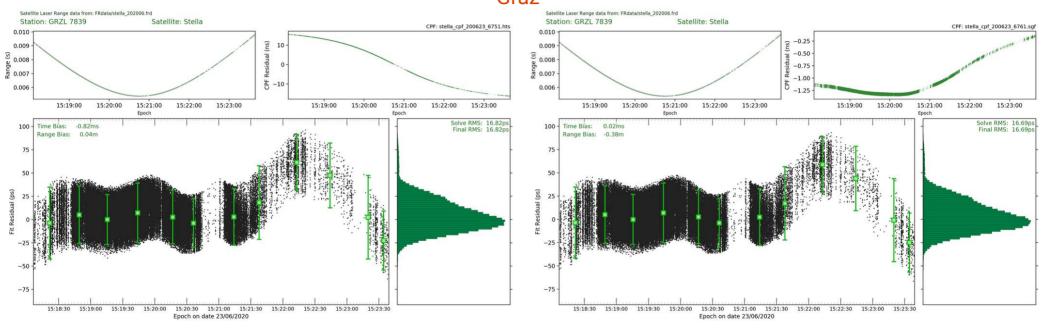


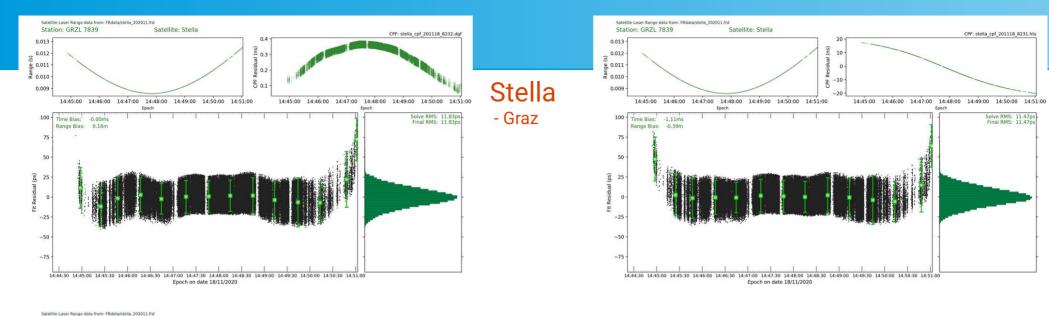


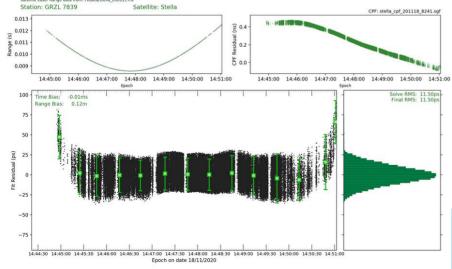




Stella - Graz











7838 Simosato Update January 2021

Van S Husson vhusson@peraton.com ILRS Quality Control Board



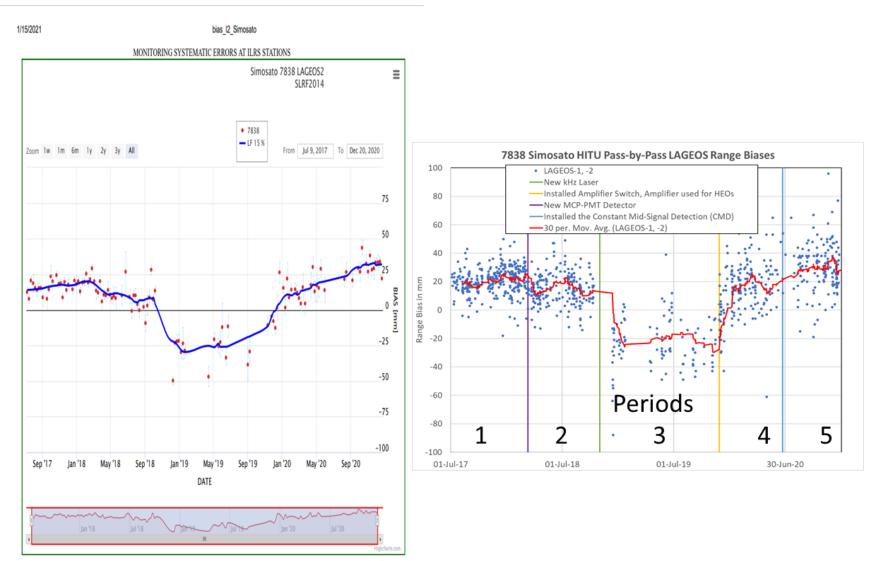
- Site Log was updated in December 2020 and approved January 2021
 - Removed the '04' occupation in Section 3
 - Updated Calibration Section 8 for clarity
 - March 1, 1982 to June 30, 2006: Calibration target at 1414.699 meters (m)
 - July 1 to August 3, 2006: experimented with 16.490 m target
 - August 4-31, 2006: remeasured long target, new range was 1414.710 m (11 mm change)
 - September 1-14 2006: experimented again with 16.490 m target
 - September 15, 2006 to June 3, 2015: long target at 1414.710 m
 - June 4, 2015 to present: target at 0 m (on the telescope frame)

Station History

- > No change since last meeting, they asked advice on the data impact flag
- > They still need to merge their old history log with their new history log



JCET and HITU LAGEOS 7838 Range Bias Analysis Peraton



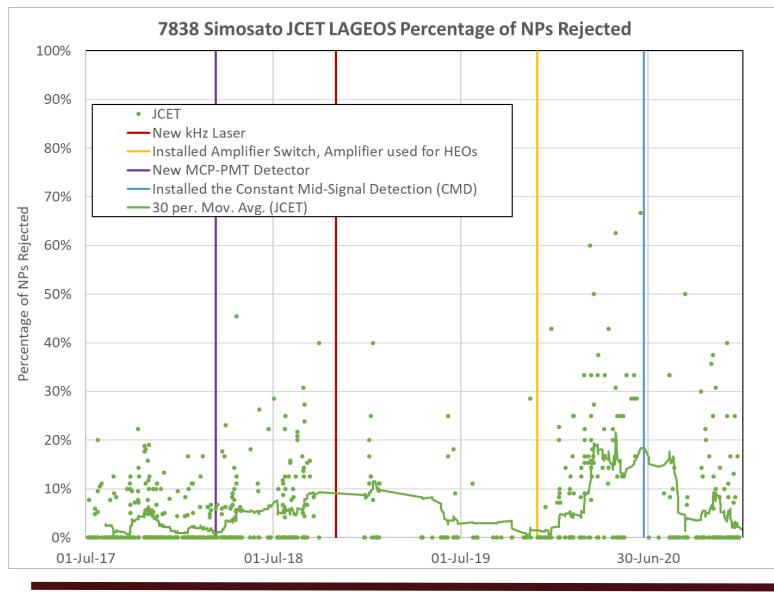
JCET weekly L2 biases on the left And HITU L1 and L2 pass-by-pass biases on the right. Good agreement after editing gross outliers:

Gross LAGEOS Outliers since July 2017: June 1 to 4, 2018: 180 μsec time bias (see Rapidmail # 137)

Following L2 passes had meter biases: April 7, 2020 at 2:48 April 8, 2020 at 0:45 and 4:49 August 5, 2020 at 6:01



7838 JCET NP Rejection Percentage by Pass Peraton



In the daily bias reports from the Analysis Centers, range and time biases are computed per pass along with the number of normal points (NPs) accepted and rejected. HITU usually accepts all normal points including Simosato, but the other centers have higher NP rejection rates.

On the left is the JCET pass-by-pass LAGEO NP rejection rates for each Simosato pass.

Are these inconsistent NPs within a pass kept in the weekly solution estimates?

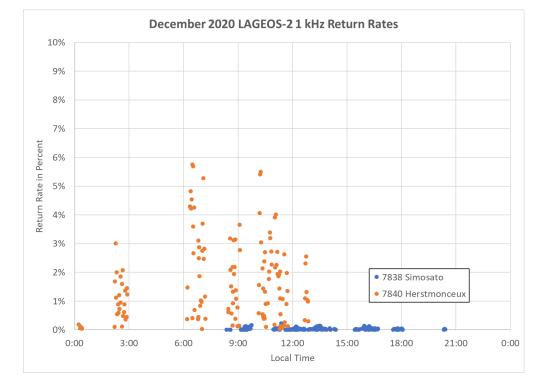
If they are kept are they corrupting the solution?



Dec 2020 1kHz LAGEOS-2 Return Rate Analysis Peraton

Herstmonceux Laser

Laser Type : ND:YAG Secondary Wavelength [nm]: 532 Secondary Max. Energy [mJ]: 1.0 Pulse Width (FWHM) [ps]: 10 Max. Repetition Rate [Hz]: 1000 Fullw. Beam Divergence ["]: 5 - 200 Final Beam Diameter [m]: 0.03



Simosato Laser

Laser Type : ND:YAG Secondary Wavelength [nm]: 532 Secondary Max. Energy [mJ]: 3 Pulse Width (FWHM) [ps]: 30 Max. Repetition Rate [Hz]: 1,000 Fullw. Beam Divergence ["]: 6 - 20 Final Beam Diameter [m]: 0.75

Simosato operated mostly in the day time in December. Simosato maximum LAGEOS -2 receive rate in December 2020 was 0.2 %, and their average return rate was 0.03 %. While Herstmonceux average return rate was 1.7 %.