



Joint-stock Company  
«RESEARCH-AND-PRODUCTION CORPORATION  
«PRECISION SYSTEMS AND INSTRUMENTS»»

# **Preliminary results of the ILRS network performance in the LARGE-3 experiment**

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**Matera, 2015**



# Targets and definitions of the LARGE-3 experiment

## Targets:

**GLONASS-123, -125, -128, -129, -133, -134**

**Compass-M3**

**Galileo-101, -102, -103 and -104**

## Definitions:

**NP length:** 5 minutes or 1000 FR points.

**Sector definition:** The duration of each satellite's visibility is divided into 3-sectors (beginning, middle, and end).

**Pass definition:** A pass is counted for each visibility where a station tracked the satellite.



# LARGE-3 requirements

**This time we need (Third GNSS SLR Tracking Campaign):**

**1. 9 normal points per one pass from each station:**

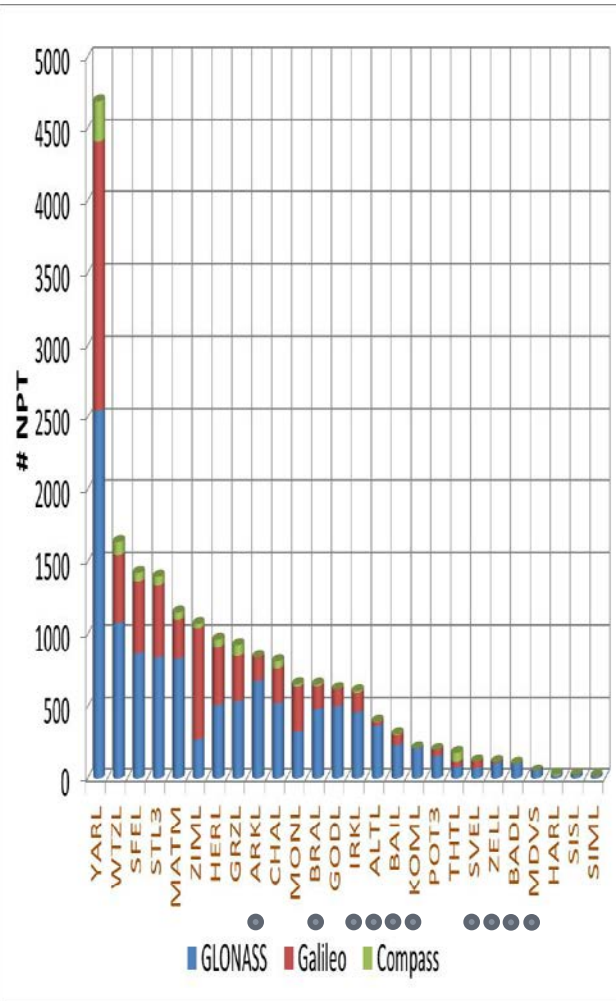
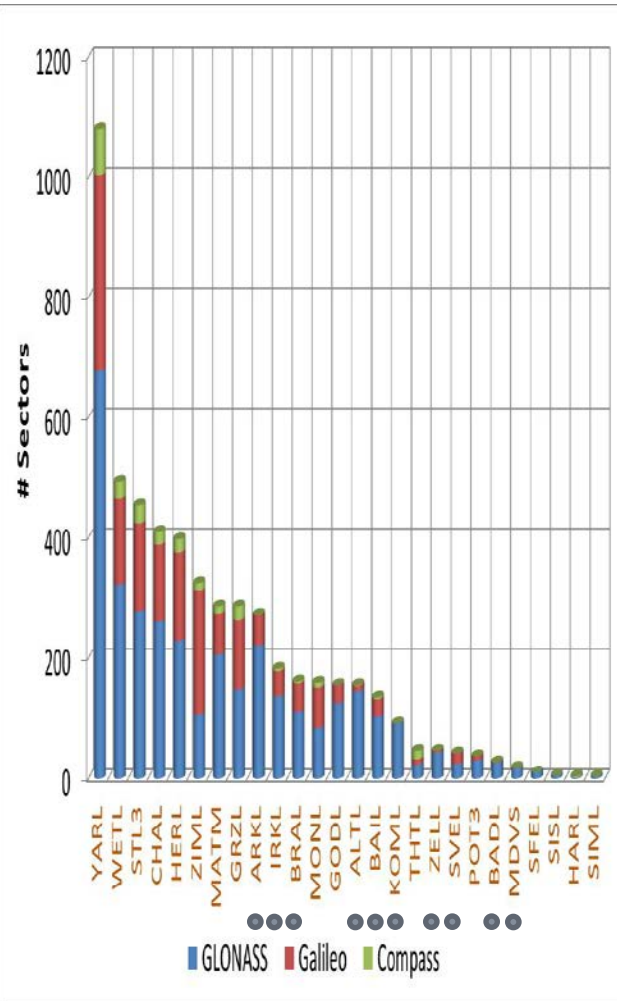
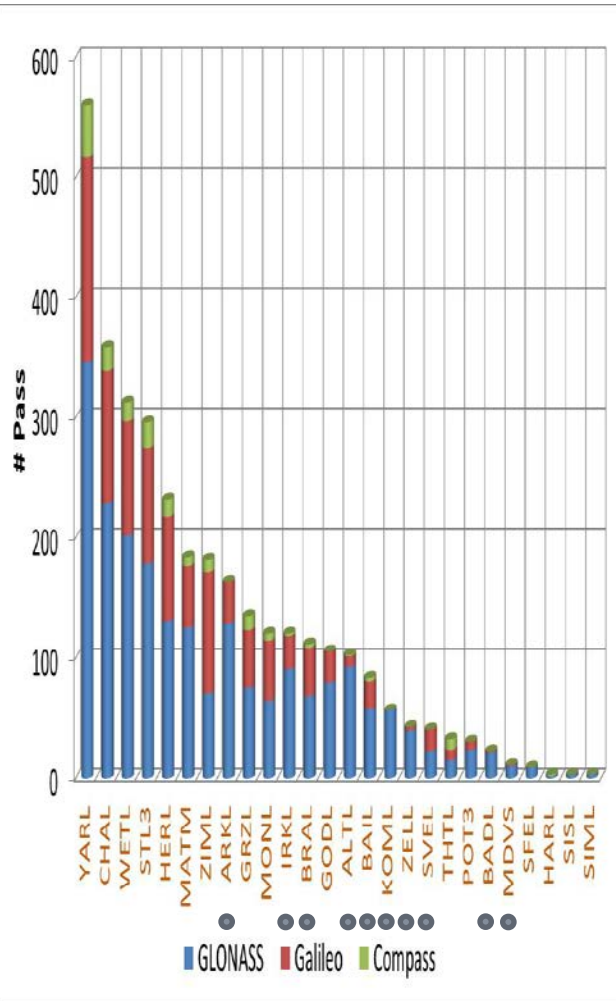
- **3 points in an ascending section (at the beginning of the pass),**
- **3 points in the middle section of the pass, about the traverse point, and**
- **3 points in the descending section (at the end of the pass).**

**In each section NPs may be taken together or separately whatever is better for your operation.**

**2. More daytime ranging even if it is around sunrise and sunset.**



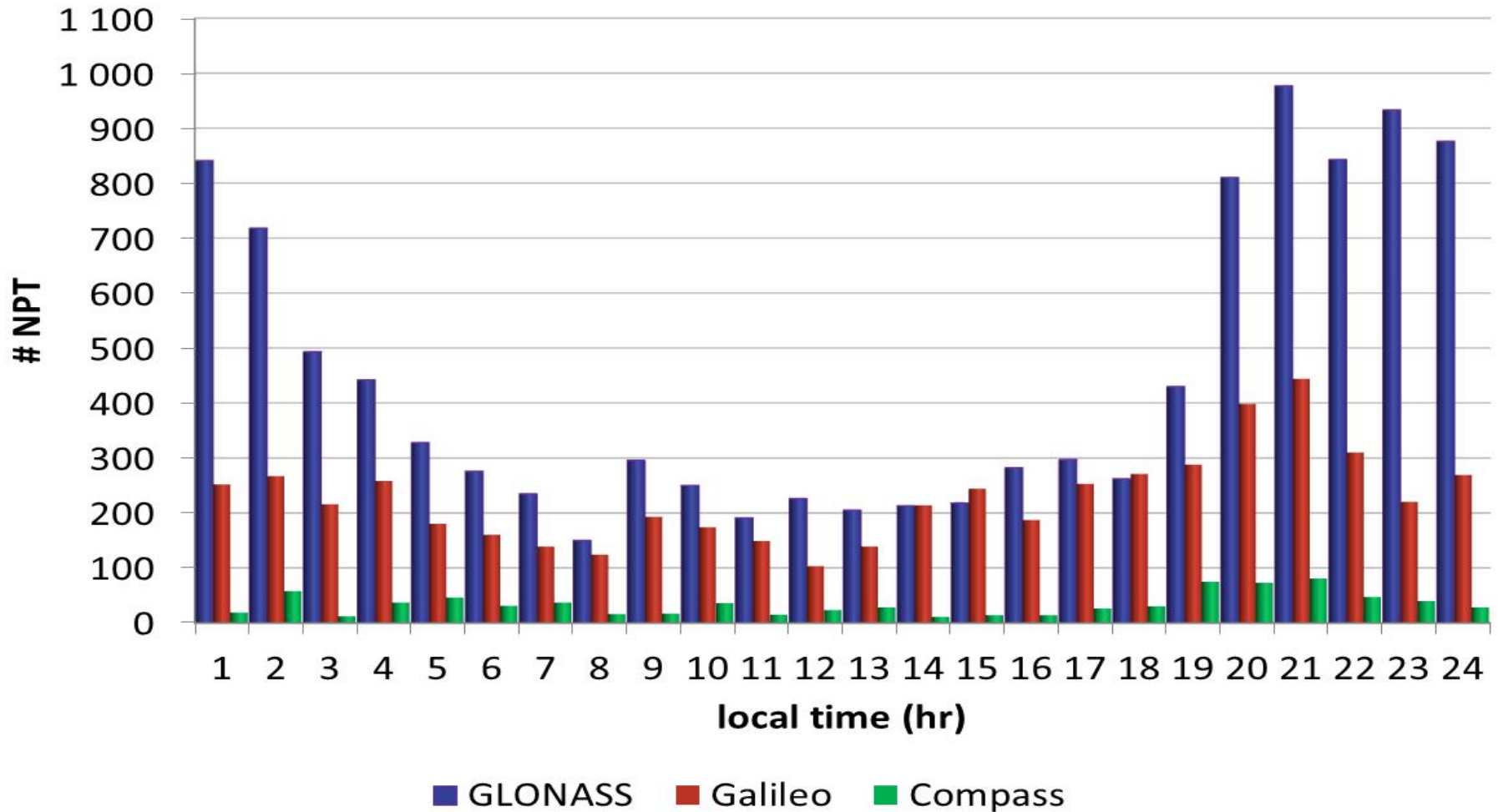
# Passes / Sectors / Normal points



**Conclusion:** Required (expected) ratio  $P/S/NP = 1/3/9$ . Actual (LARGE-3): min=1/1.1/2.3 (CHAL), max=1/1.9/8.4 (YARL).



# Dependency between a number of sessions and the local time



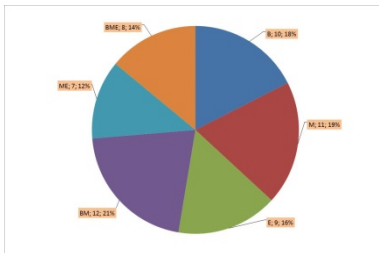
**Conclusion:** the majority of SLR observations complies with night conditions of functioning, maximum measurements taken at 9 p. m. local time



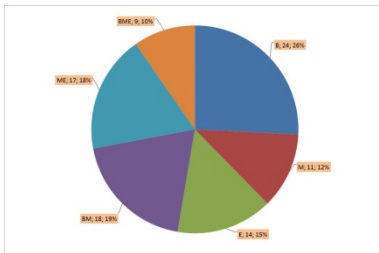
# Typical distribution of NP on sectors and passes

Type 1 («quasi-uniform»)

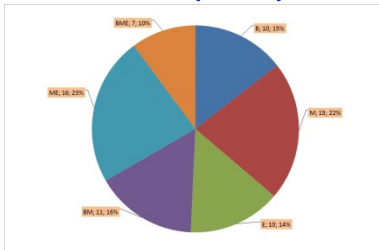
1868 (KOML)



1879 (ALTL)

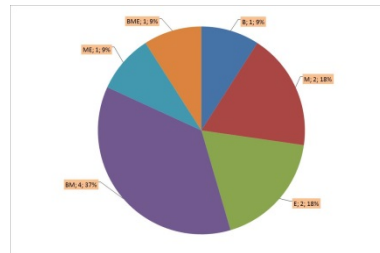


7407 (BRAL)

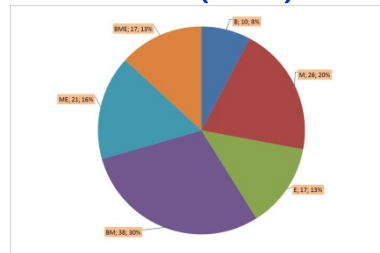


Type 2 (predominating in the ascending and middle sections)

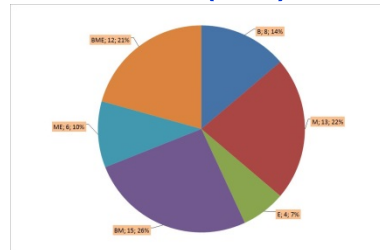
1874 (MDVS)



1886 (ARKL)

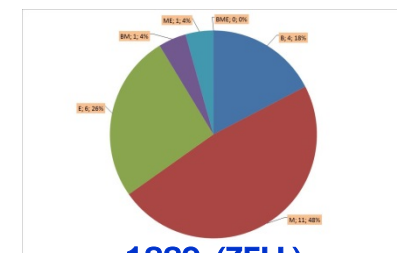


1887 (BAIL)

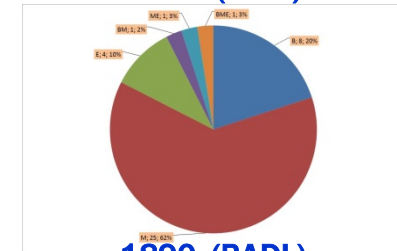


Type 3 (predominating in the middle section)

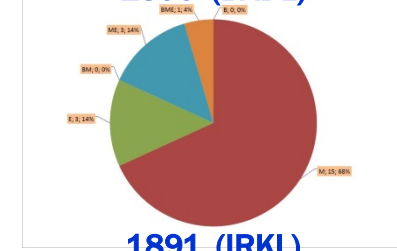
1888 (SVEL)



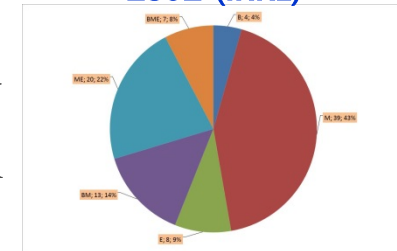
1889 (ZELL)



1890 (BADL)



1891 (IRKL)

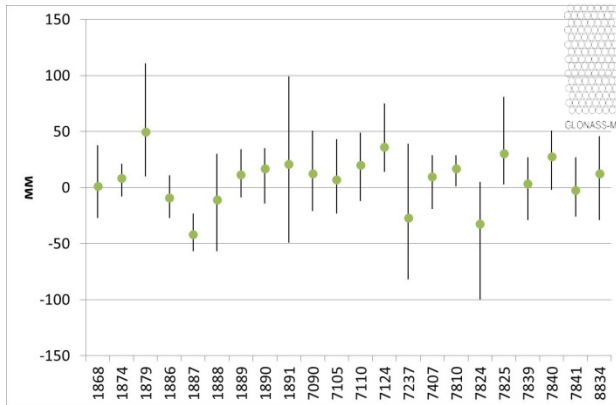


**Conclusion:** in LARGE-3, measurements from the RLRN stations do not fully comply with the requirements of uniform distribution of 9 NP per each pass.

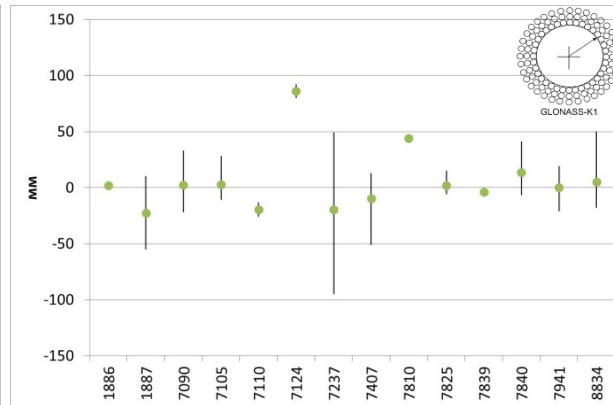


# Residuals of laser measurements on GLONASS SC in relation to the «laser» orbit

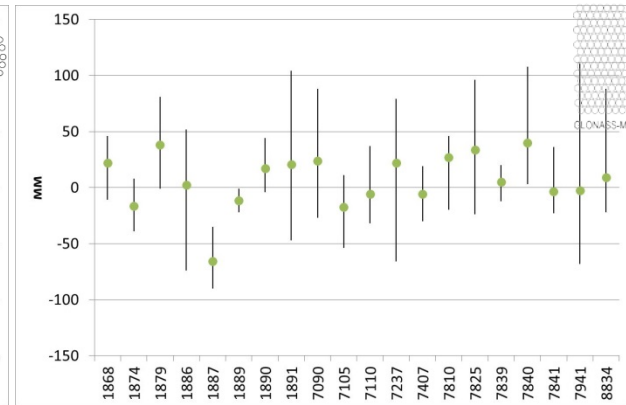
## GLONASS-123



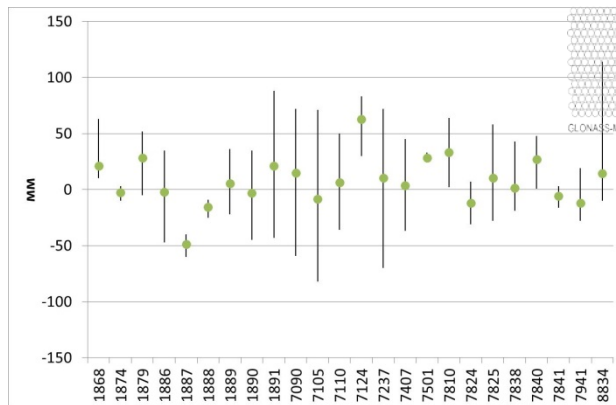
## GLONASS-125



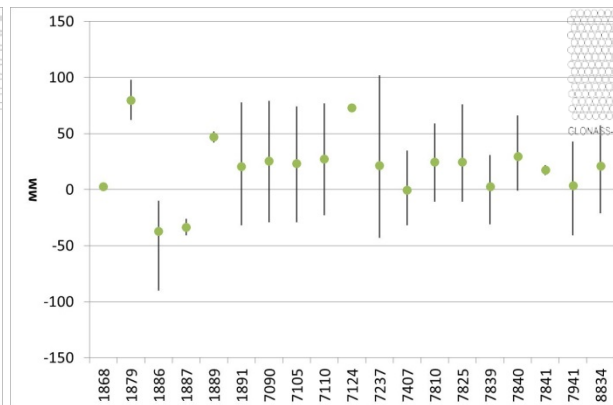
## GLONASS-128



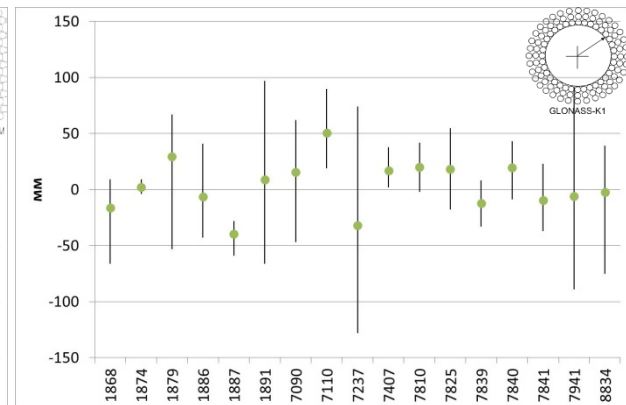
## GLONASS-129



## GLONASS-133



## GLONASS-134

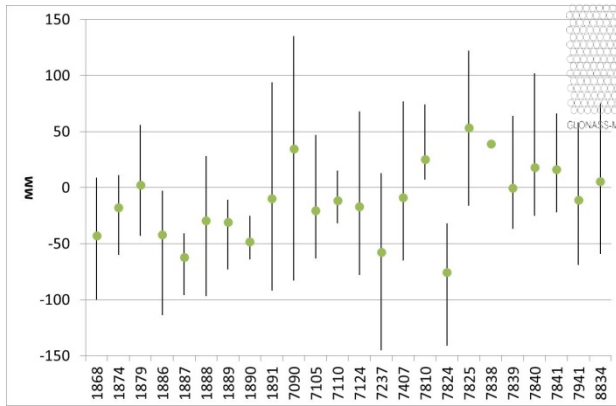


**Conclusion:** Laser measurements on GLONASS SC are characterized by spread of residual average values for different means varying from **-65.8** to **86.0** mm with  $\sigma \leq 33.7$  mm. Given that the probability equals 0.95, the residual average value deviation for all the means does not exceed **48.7** mm.



# Residuals of laser measurements on GLONASS SC in relation to the «radio» orbit

### GLONASS-123

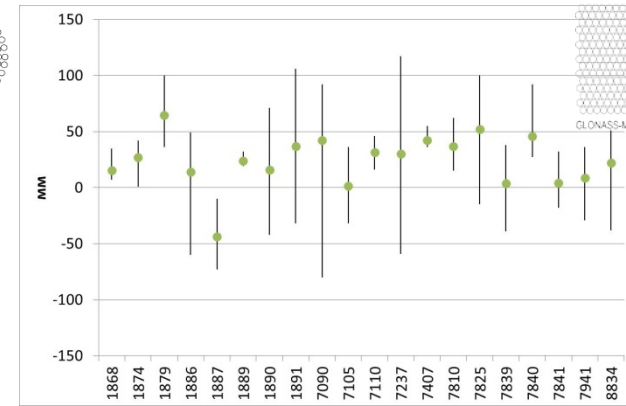


### GLONASS-125

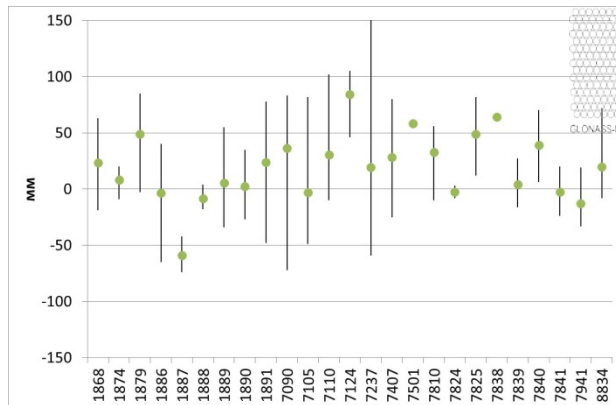


«Radio» orbit is N/A

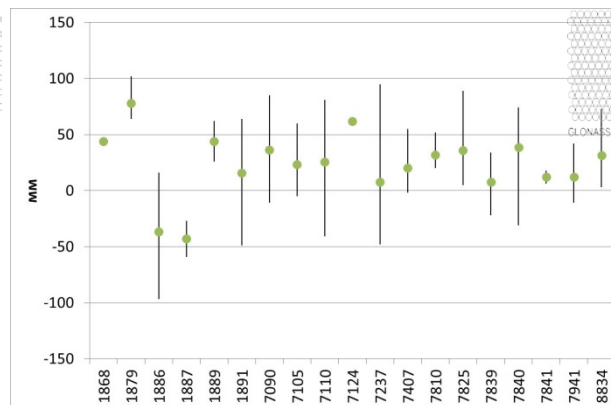
### GLONASS-128



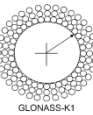
### GLONASS-129



### GLONASS-133



### GLONASS-134



«Radio» orbit is N/A

**Conclusion:** laser measurements on GLONASS SC are characterized by spread of residual average values for different means varying from **-75.5** to **84.3** mm with  $\sigma \leq 36.3$  mm. Given that the probability equals 0.95, the residual average value deviation for all the means does not exceed **62.0** mm.





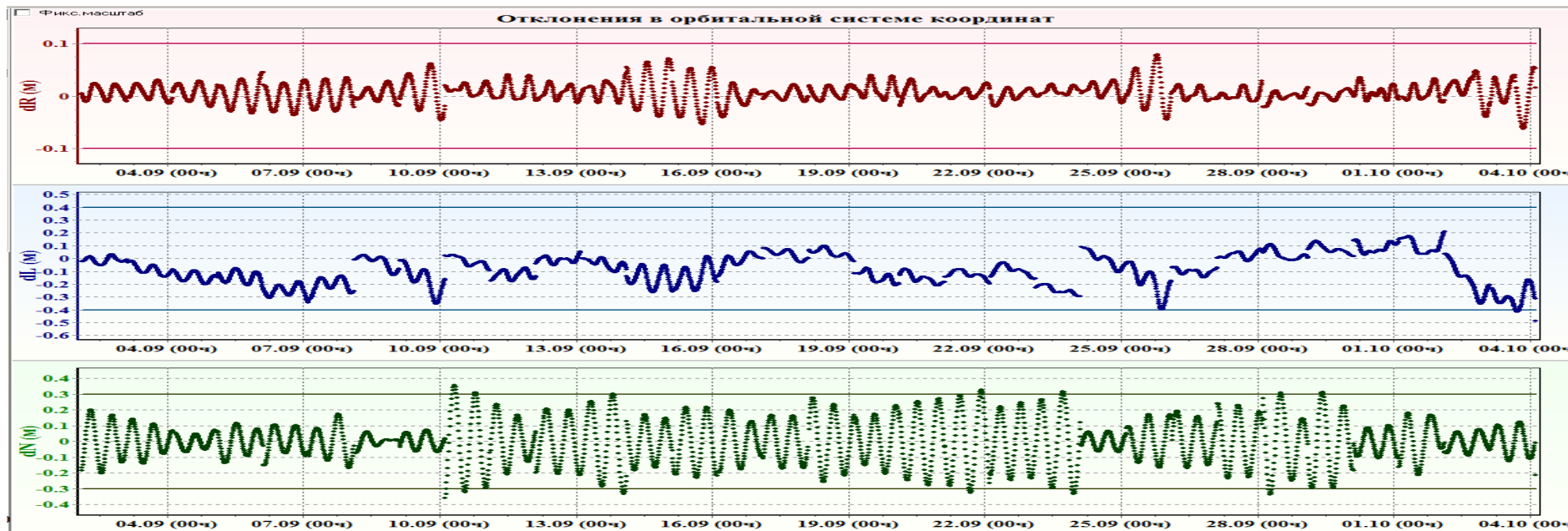
# Accuracy of GLONASS SC «laser» orbits

Measuring interval (MI): 4 days.

Number of laser measurements on the MI (LARGE-3): from 25 to 44 sessions.

«Laser» orbit: average (2nd and 3rd) days of the MI.

## GLONASS-129



Note: «Radio» orbit – SVOEVP / SHPDETC ([www.glonass-svoevp.ru](http://www.glonass-svoevp.ru))

**Conclusion:** difference between the GLONASS SC «radio» and «laser» orbits is characterized by  $\sigma_r \sim 1.8$  cm,  $\sigma_l \sim 11.2$  cm,  $\sigma_n \sim 14.2$  cm.



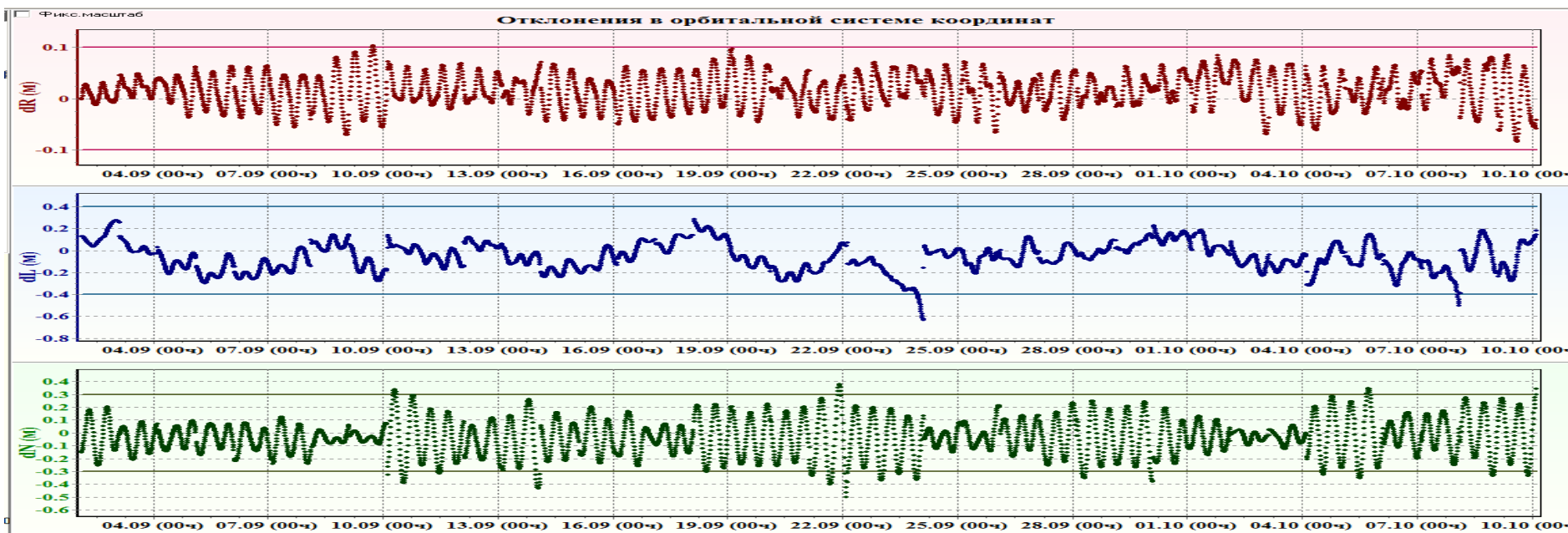
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Number of laser measurements on the MI (LARGE-3): from 25 to 44 sessions.

«Laser» orbit: average (2nd and 3rd) days of the MI.

## GLONASS-129



Note: «Radio» orbit – IAC TSNIIMASH ([www.glonass-iac.ru](http://www.glonass-iac.ru))

**Conclusion:** difference between the GLONASS SC «radio» and «laser» orbits is characterized by  $\sigma_r \sim 3.2$  cm,  $\sigma_l \sim 12.4$  cm,  $\sigma_n \sim 14.3$  cm



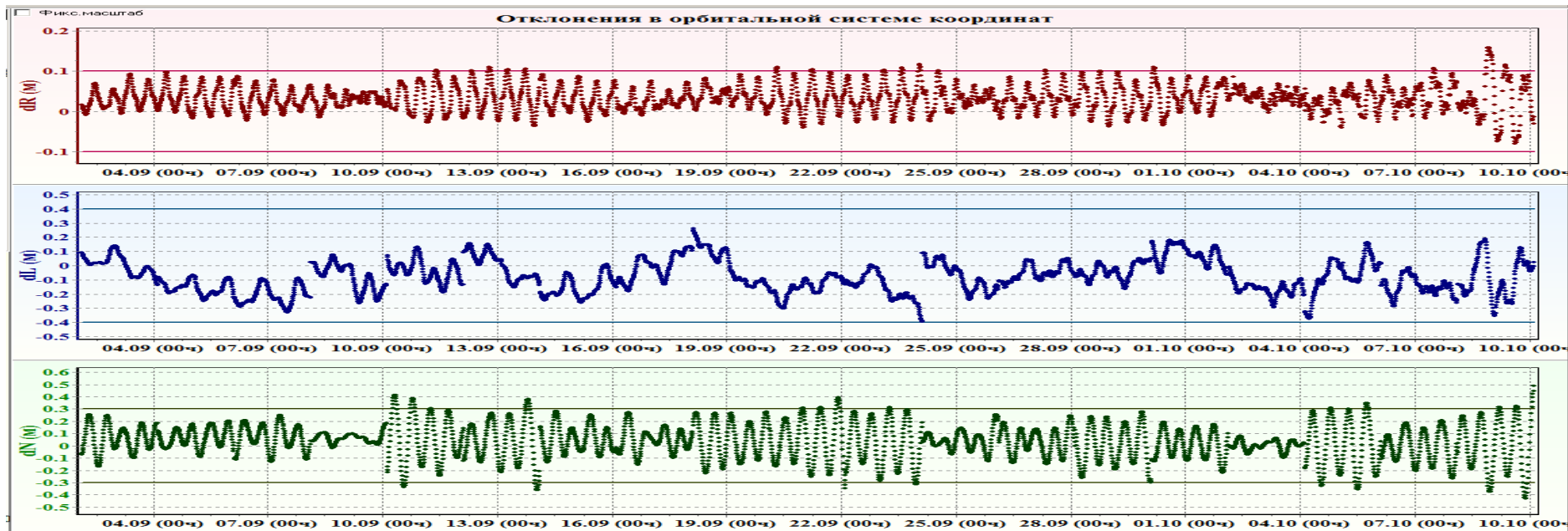
# Accuracy of GLONASS SC «laser» orbits

Measuring interval (MI): 4 days.

Number of laser measurements on the MI (LARGE-3): from 25 to 44 sessions.

«Laser» orbit: average (2nd and 3rd) days of the MI.

## GLONASS-129



Note: «Radio» orbit – AC CODE ([www.aiub.unibe.ch](http://www.aiub.unibe.ch))

**Conclusion:** difference between the GLONASS SC «radio» and «laser» orbits is characterized by  $\sigma_r \sim 3.1$  cm,  $\sigma_l \sim 11.2$  cm,  $\sigma_n \sim 14.0$  cm



## LARGE-3 experiment generalizations (positive)

- 1. Within the framework of LARGE-3, GLONASS SC have been tracked by 23 ILRS and 10 RLRN stations.**
- 2. We have taken measurement on all planned targets.**
- 3. The performance rate of stations has significantly increased on the interval of the experiment duration – it was up to 15-20 normal points from the ILRS network.**
- 4. It is the first time the LARGE-3 experiment has provided an opportunity to determine SC «laser» orbits more accurately than the IGS orbits, but to achieve a theoretically possible accuracy it is necessary to increase the amount of measurements taken at least two times.**



## LARGE-3 experiment generalizations (negative)

- 1. SLR observations are non-uniformly distributed both on SC orbit passes and between Northern and Southern hemispheres of the Earth.**
- 2. The number of measurement taken at night (local astronomical time) significantly (20 and more times) exceeds the number of daytime measurements.**
- 3. In some cases, there are no SLR observations on some passes at all.**



## Propositions on further SLR development

- 1. To increase the number of daytime measurements. Examples – stations in Graz, Wettzell, Yarragadee and Brasilia.**
- 2. When planning SLR observations, it is necessary to provide a more uniform coverage of GNSS orbit arcs by measurements taken using the stations in the Northern and Southern hemispheres and on each pass.**
- 3. When developing the new generation millimeter accuracy SLR stations, it is required to increase their production rates with regard to GNSS SC to 12-24 NP per hour.**

**Our next report called «Satellite radio laser ranging stations for application in GNSS: requirements for the technical characteristics and methods of their implementation» is dedicated to study on how to significantly increase the number of SLR observations on each GNSS SC.**



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